

Many-body forces and nucleon clustering near the QCD critical point

- *E. Shuryak and J. M. Torres-Rincon, Phys. Rev. C100, 024903 (2019), arXiv:1805.04444 [hep-ph].*
- *E. Shuryak and J. M. Torres-Rincon, Phys. Rev. C101, 034914 (2020),*
- *arXiv:1910.08119 [nucl-th].*
- *D. DeMartini and E. Shuryak, (2020), arXiv:2007.04863 [nucl-th].*

Many-body forces and nucleon clustering near the QCD critical point D. DeMartini and E. Shuryak, e-Print: 2010.02785

1. Introduction: suggestion of BES, critical event-by-event fluctuations
2. The main idea: preclusters may have **size comparable to corr.length**
3. **Nucleon clustering, importance of 4 N systems, kurtosis and viral expansion**
4. Paradox: attractive binary forces near CP get huge $O(N^2)$
5. **Repulsive** manybody forces near CP, estimates in Landau model
6. **universal effective action** for Ising universality class
7. **Deformed** universal effective action
8. Summing all effects near CP
9. **EXPERIMENTAL OBSERVABLES: Kurtosis and t^*p/d^2 plots: where CP MAY be located?**

Introduction

- **The original ideas:**
- **Look for event-by-event fluctuations**
- **Perform beam energy scan**
- **Watch for non-monotonous signals**

M. A. Stephanov, K. Rajagopal, and E. V. Shuryak,
Phys. Rev. Lett. 81, 4816 (1998), arXiv:hep-ph/9806219 [hep-ph].

Higher moments of the critical field

$$\kappa_2 = \langle \phi^2 \rangle, \quad \kappa_3 = \langle \phi^3 \rangle, \quad \kappa_4 = \langle \phi^4 \rangle - 3\langle \phi^2 \rangle^2$$

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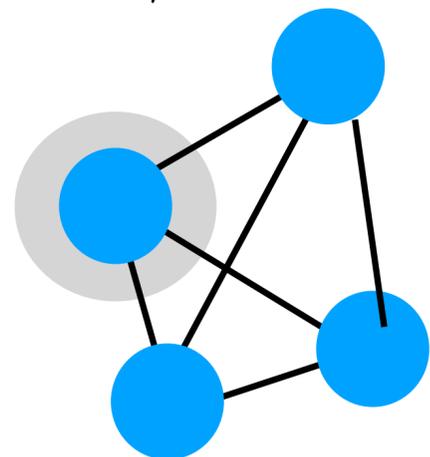
Yet one cannot directly measure moments of ϕ ...
they are related to moments of nucleon multiplicity distribution, but **not trivially**

So far estimates relied on the assumption that nucleons
are correlated **ONLY** due to near-CP fluctuations,
which is of course **not the case**

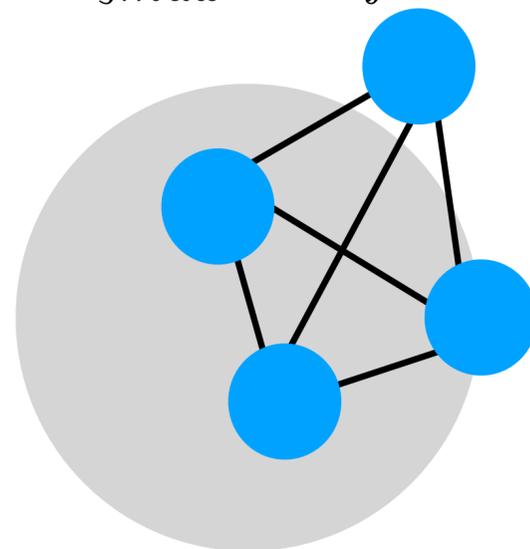
The main idea of this work:

Suppose the CP indeed exists, and is located in the part of the phase diagram near the freezeout line of BES program collisions. Furthermore, while scanning this line, for some specific beam energy one happens to be in a state in which the correlation length reaches a value $\xi_{max} \sim 1.5-2\text{fm}$. What observables are sensitive to such scale of ξ ?

$$\xi_{min} \sim 1/m_\sigma \sim 0.4 \text{ fm}$$



$$\xi_{max} \sim 2 \text{ fm}$$



Pre-clustering of nucleons create objects of the right scale !

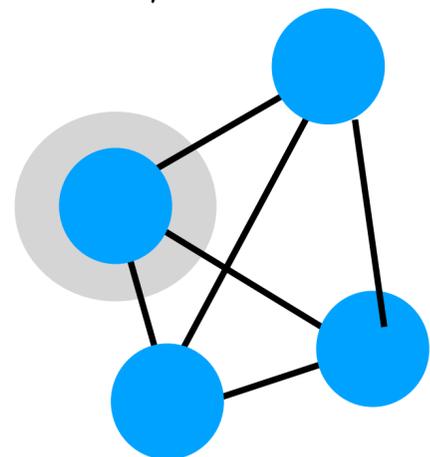
Their energy — and therefore production yield — is very sensitive to correlation length

As we will show, the interplay of attractive binary
And repulsive manybody forces
Will lead to non-monotonous signal

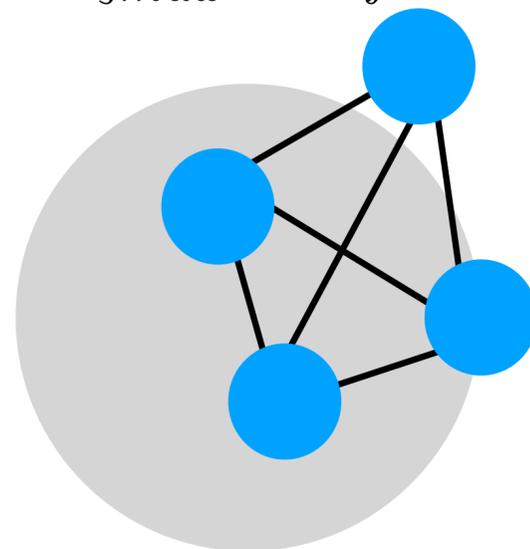
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Side remark: sound waves we observed,
but with the wavelength much larger than 2 fm

Studies of few-nucleon pre-clustering at freezeout conditions

Classical molecular dynamics

- *E. Shuryak and J. M. Torres-Rincon, Phys. Rev. C100, 024903 (2019), arXiv:1805.04444 [hep-ph].*

Semiclassical approximation
(fluctons)
At finite temperatures

- *E. Shuryak and J. M. Torres-Rincon, Phys. Rev. C101, 034914 (2020),*
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K-harmonics => radial Schrodinger
equations in $3*(N-1)$ dimensions

Direct Path Integral
Monte Carlo (PIMC)
Numerical simulations

- *D. DeMartini and E. Shuryak, (2020), arXiv:2007.04863 [nucl-th].*

(the first time ever) testing the flucton method at finite T

Fluctons for anharmonic oscillator at $T \neq 0$

$$S_E = \oint d\tau \left(\frac{\dot{x}^2}{2} + \frac{x^2}{2} + \frac{g}{2}x^4 \right) .$$

the usual density matrix (line, 60 states)

$$P(x_0) = \sum_i |\psi_i(x_0)|^2 e^{-E_i/T}$$

$$P(x_0) \sim \exp\left(-S_E[x_{flucton}(\tau)]\right)$$

(points on the plot)
so, the method works very well

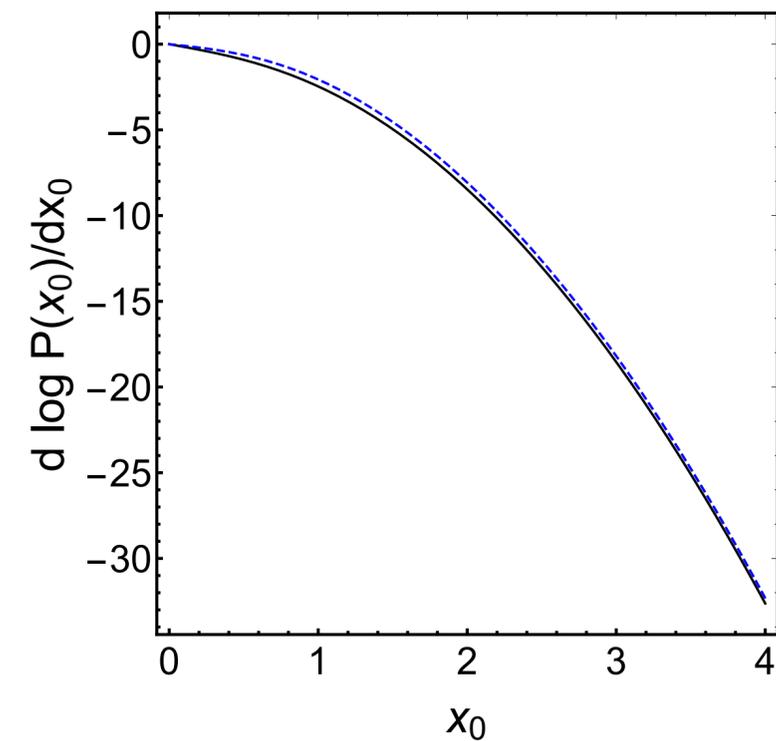
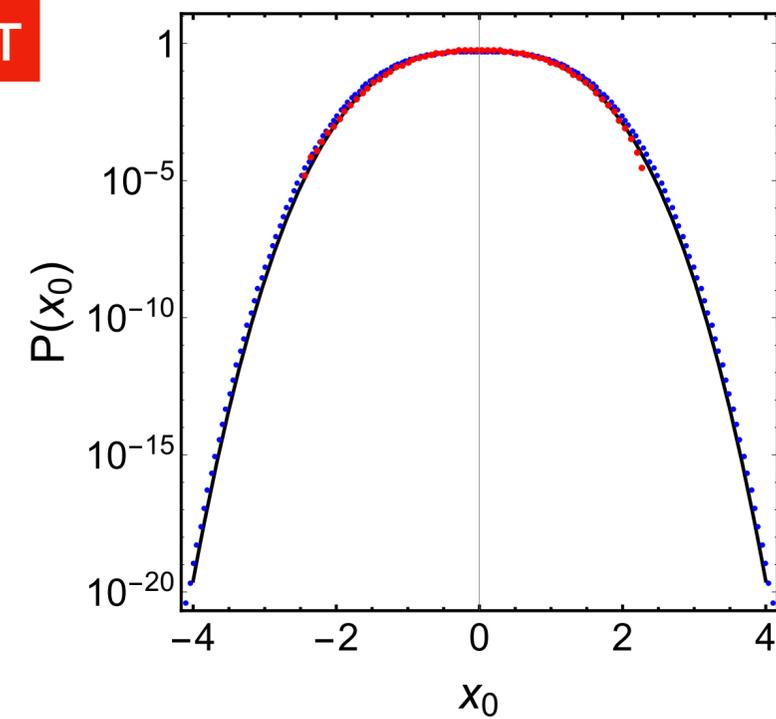


FIG. 3: Top panel: Density matrix $P(x_0)$ vs x_0 for anharmonic oscillator with the coupling $g = 1$, at temperature $T = 1$, calculated via the definition (1) (line) and the flucton method (points). The line is based on 60 lowest state wave functions found numerically. Bottom panel: Comparison of the logarithmic derivative of the density matrix of the upper panel.

**K-harmonics applied to He4 (not a new method,
and yet we found something new with it...)**

9 Jacobi coordinates for 4 particles

$$\vec{\xi}[1] = \frac{\vec{x}[1] - \vec{x}[2]}{\sqrt{2}}, \quad \vec{\xi}[2] = \frac{\vec{x}[1] + \vec{x}[2] - 2\vec{x}[3]}{\sqrt{6}},$$

hyperdistance

**in 9 dimensional space
Is sum of squares of all 6
Distances**

$$\vec{\xi}[3] = \frac{\vec{x}[1] + \vec{x}[2] + \vec{x}[3] - 3\vec{x}[4]}{2\sqrt{3}}$$

$$\rho^2 = \sum_{m=1}^3 \vec{\xi}[m]^2 = \frac{1}{4} \left(\sum_{i \neq j} (\vec{x}[i] - \vec{x}[j])^2 \right)$$

**redefining the wave function
and the radial Schreodinger eqn
Note, the first derivative is gone
but **some new repulsive
potential remains (not orbital!)****

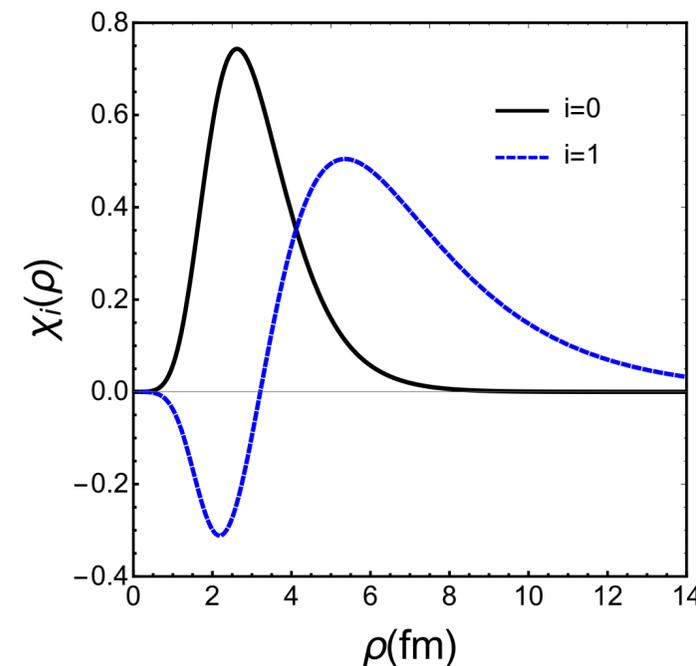
$$\psi(\rho) = \chi(\rho) / \rho^4$$

$$\frac{d^2 \chi}{d\rho^2} - \frac{12}{\rho^2} \chi - \frac{2M}{\hbar^2} (W(\rho) + V_C(\rho) - E) \chi = 0$$

Solving the eigenvalue problem in App. A we have obtained 40 lowest eigenstates for Eq. (A3) using the simplest potential V_1 from Ref. [17] and the Coulomb term between the two protons. The ground state energy we find is $E_0 = -27.8$ MeV, very close to the experimental value of -28.3 MeV.

Rather unexpectedly, we also find a second bound state (missed in [17]) with energy $E_1 = -2.8$ MeV. To determine whether this state is physical, we show in Table ?? the excited states of ${}^4\text{He}$. Among them there is just one 0^+ state, with a binding energy of

$$B = -28.3 \text{ MeV} + 20.2 \text{ MeV} = -8.1 \text{ MeV}$$



here are experimentally observed
excited states of He4
the first one fits well
to our second bound state

Now, getting convinced
that we understand
quantum mechanics of 4 nucleons in He4
At zero T,
we proceed to calculate
the density matrix at finite T
and check how it changes
when the nuclear potential changes

So, people doing stat models
For light nuclei
Were missing about 50 states!

TABLE I: Low-lying resonances of ${}^4\text{He}$ system, from BNL properties of nuclides listed in *nndc.bnl.gov* web page. J^P is total angular momentum and parity, Γ is the width. The last column is the decay channel branching ratios, in percents. p, n, d correspond to emission of proton, neutron or deuterons.

E (MeV)	J^P	Γ (MeV)	decay modes, in %
20.21	0^+	0.50	p =100
21.01	0^-	0.84	n =24, p =76
21.84	2^-	2.01	n = 37, p = 63
23.33	2^-	5.01	n = 47, p = 53
23.64	1^-	6.20	n = 45, p = 55
24.25	1^-	6.10	n = 47, p = 50 , d=3
25.28	0^-	7.97	n = 48 , p = 52
25.95	1^-	12.66	n = 48 ,p = 52
27.42	2^+	8.69	n = 3 , p = 3 ,d = 94
28.31	1^+	9.89	n = 47 , p = 48 , d = 5
28.37	1^-	3.92	n = 2, p = 2, d = 96
28.39	2^-	8.75	n = 0.2, p = 0.2 , d = 99.6
28.64	0^-	4.89	d=100
28.67	2^+	3.78	d=100
29.89	2^+	9.72	n = 0.4 , p = 0.4, d = 99.2

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Now they are being included: About half of d,t,he3
Come from them!

I did PIMC simulation of He4 in 1980 already, and managed to put it to NPB

Path integral simulations of the few-nucleon clustering at heavy ion collisions freezeout,

(with Dallas DeMartini, SB student)

hyperdistance definition

$$\rho^2 = \sum_{m=1}^3 \xi[\vec{m}]^2 = \frac{1}{4} \left(\sum_{i \neq j} (\vec{x}[i] - \vec{x}[j])^2 \right)$$

4 particles, 6 distances

Paths of 4 nucleons
in a Matsubara time

In a periodic box

Only tau discretized

But very many steps needed

The density and
temperature values
correspond to kinetic
Freezeout conditions
at BES 1 energies

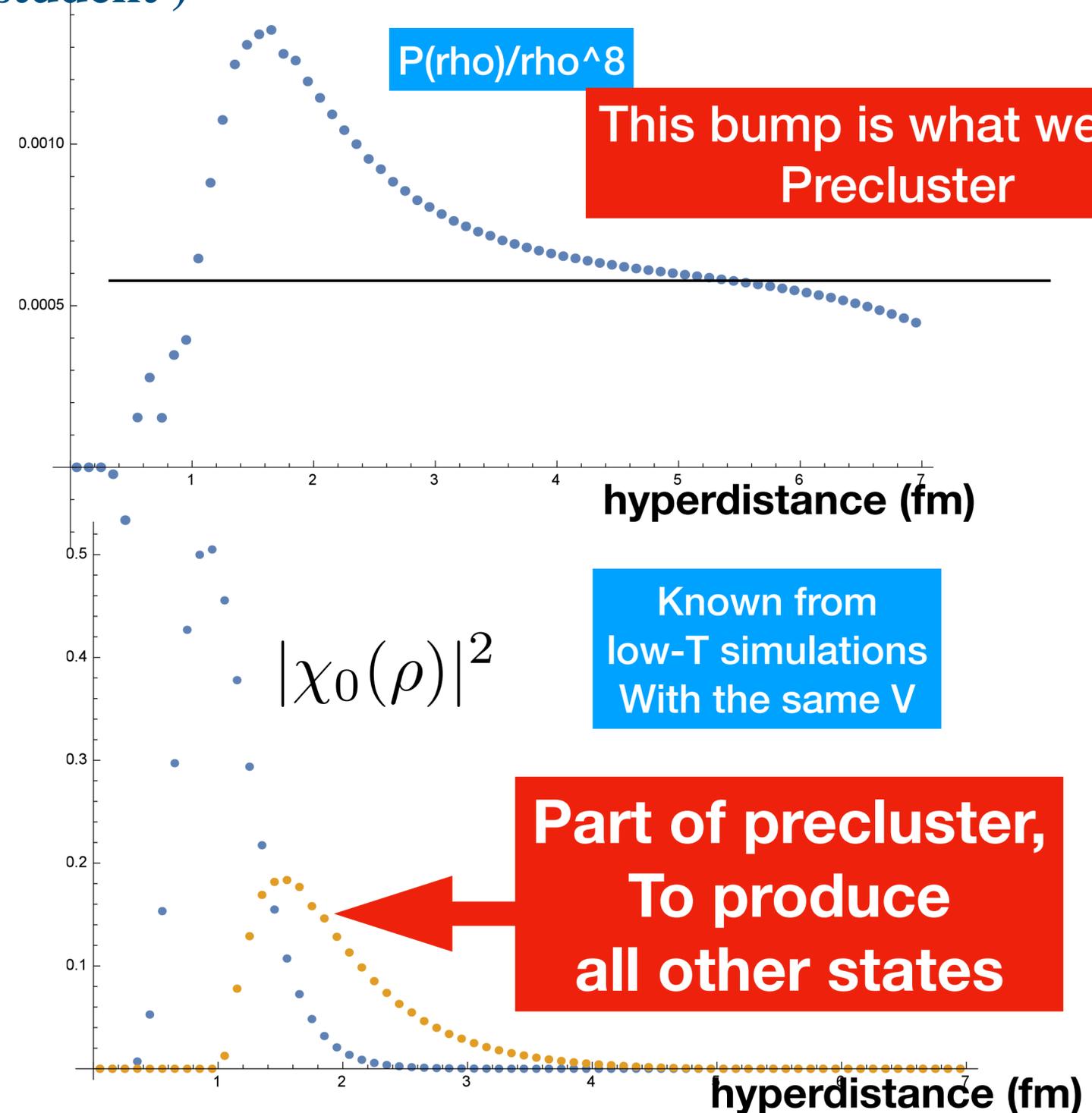


FIG. 6: Ground state and precluster for 19.6 GeV.

**Calculation with
Conventional nuclear forces**

$$Z_{pot} = 1 + \frac{1}{V^N} \int d^3x_1 \dots \int d^3x_N [e^{(-\sum_{i>j} V(\vec{x}_i - \vec{x}_j)/T)} - 1]$$

$$Z_{pot} = 1 + \frac{N(N-1)(N-2)(N-3)}{4!} \left(\frac{V_{cor}}{V^3}\right) \approx 1 + n^3 V_{cor}^{(9)} \frac{N}{4!},$$

Or 4^4

$$V_{cor}^{(9)} = \frac{32}{105} \pi^4 \int d\rho \rho^8 (P(\rho) - 1).$$

$$V_{cor}^{(9)}(7.7) \approx 4.3 \cdot 10^4 \text{ fm}^9 \quad \text{From PIMC}$$

$$n_{cl} \equiv \frac{4}{(V_{cor}^{(9)})^{1/3}} \approx 0.114 / \text{fm}^3$$

about 3 times the density of
ambient matter $n_B(7.7) \approx 0.037/\text{fm}^3$.

$$R_{amb} \equiv n_B^{-1/3} \approx 3.0 \text{ fm}, \quad R_{cl} \equiv n_{cl}^{-1/3} \approx 2.0 \text{ fm}$$

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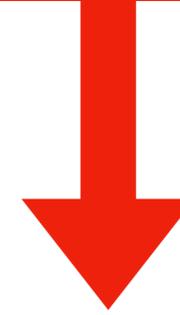
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In kurtosis it is O(1) at the lowest BES energies**



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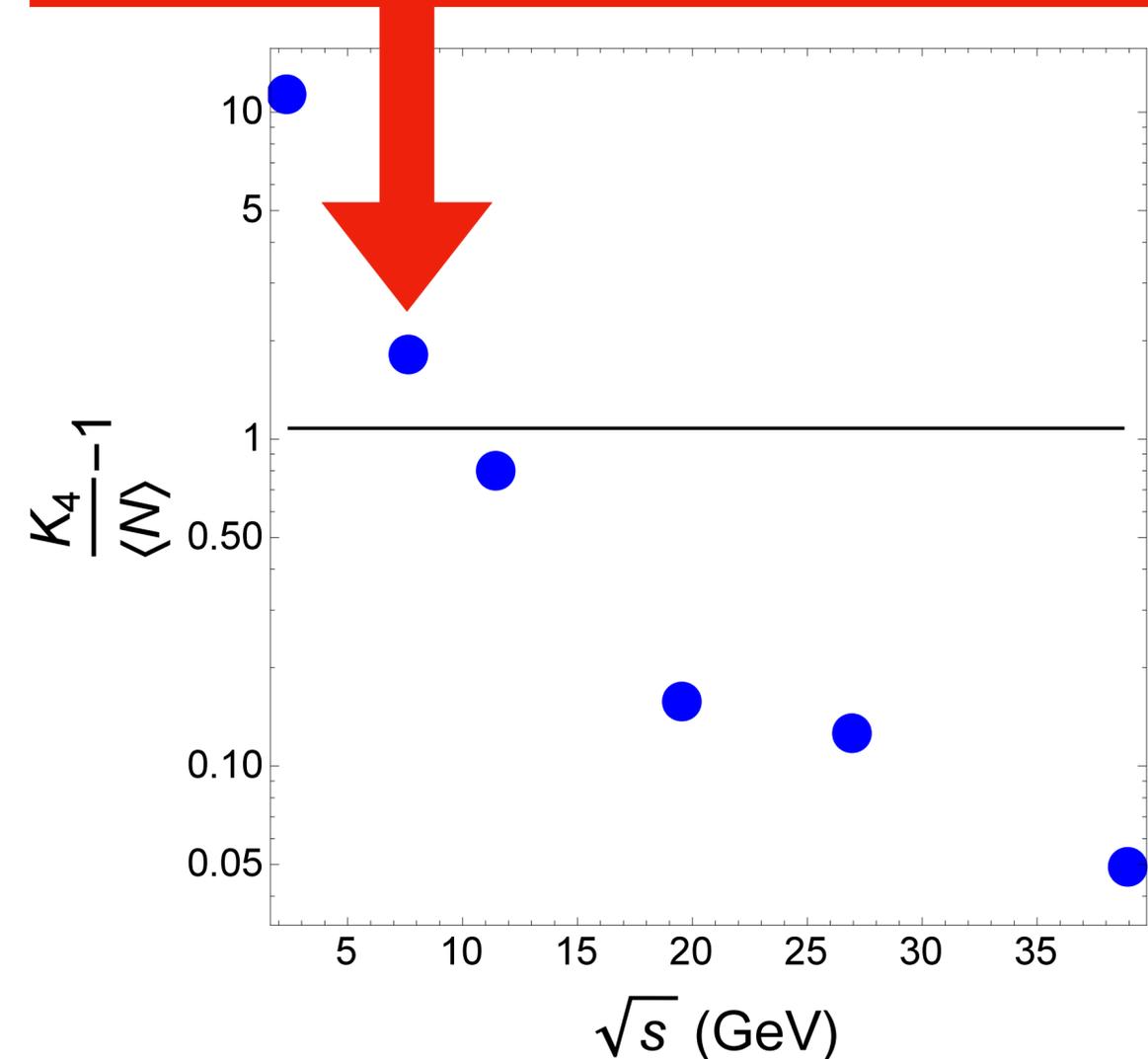


FIG. 10: The 4th cumulant deviation (Eq. (15)) versus \sqrt{s} , using the 9-dimensional correlated volume $V_{cor}^{(9)}$ determined from the PIMC simulations.

Calculation with
Modified nuclear forces
— reduced sigma mass —
Predicted by chiral transition
produce huge unrealistic effect

Predicted for
Chiral transition by RG

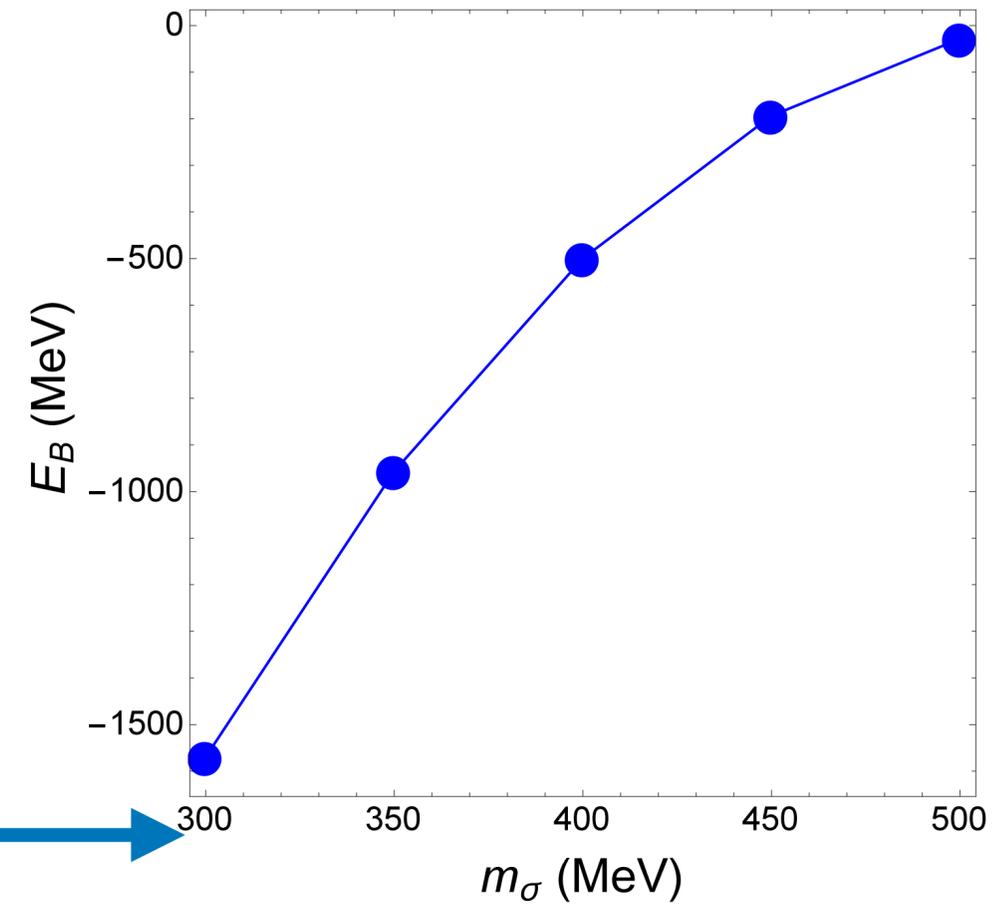
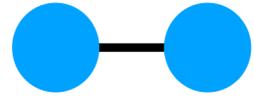


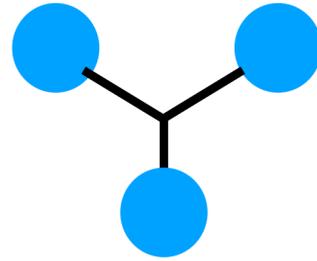
FIG. 8: Binding energy E_B of the 4N system as a function of the σ mass m_σ

the effect of binary forces induced by the critical mode at CP, where $\xi \rightarrow \infty$ must be catastrophic. Indeed, if all $N(N - 1)/2 \sim 10^4$ pairs of nucleons in the fireball be attracted to each other, with a Newton-like long-range potential, the fireball would implode, similar to a gravitational collapse.

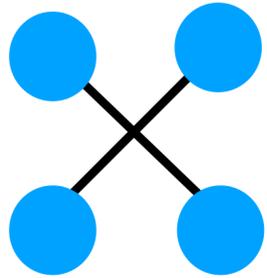
Binary and manybody forces



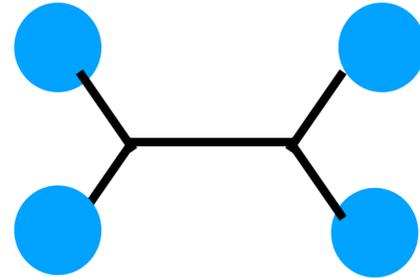
(a)



(b)



(c)



(d)

$$V_a = -\frac{g_c^2}{4\pi} \langle \phi(\vec{r}) \phi(0) \rangle = -\frac{g_c^2}{4\pi} \frac{\exp(-r/\xi)}{r}$$

We introduce the following objects

$$V_b(\vec{x}_1, \vec{x}_2, \vec{x}_3) \equiv \int d^3u D(\vec{x}_1 - \vec{u}) D(\vec{x}_2 - \vec{u}) D(\vec{x}_3 - \vec{u})$$

$$V_c(\vec{x}_1, \vec{x}_2, \vec{x}_3, \vec{x}_4) \equiv \int d^3u D(\vec{x}_1 - \vec{u}) D(\vec{x}_2 - \vec{u}) D(\vec{x}_3 - \vec{u}) D(\vec{x}_4 - \vec{u})$$

$$V_d(\vec{x}_1, \vec{x}_2, \vec{x}_3, \vec{x}_4) \equiv \int d^3u d^3v D(\vec{x}_1 - \vec{u}) D(\vec{x}_2 - \vec{u}) D(\vec{u} - \vec{v}) D(\vec{x}_3 - \vec{v}) D(\vec{x}_4 - \vec{v})$$

$$D(r) = \exp(-r/\xi)/r$$

the factor $1/4\pi$ present in 3d propagator will be included later with the couplings.

If nucleons are uncorrelated
They are easy to calculate.
But they are correlated!

These functions depend on 3 or 4 points
should be averaged over manybody
density matrix of the clusters.

How diagrams depend on the **cluster shape** and the **correlation length**?

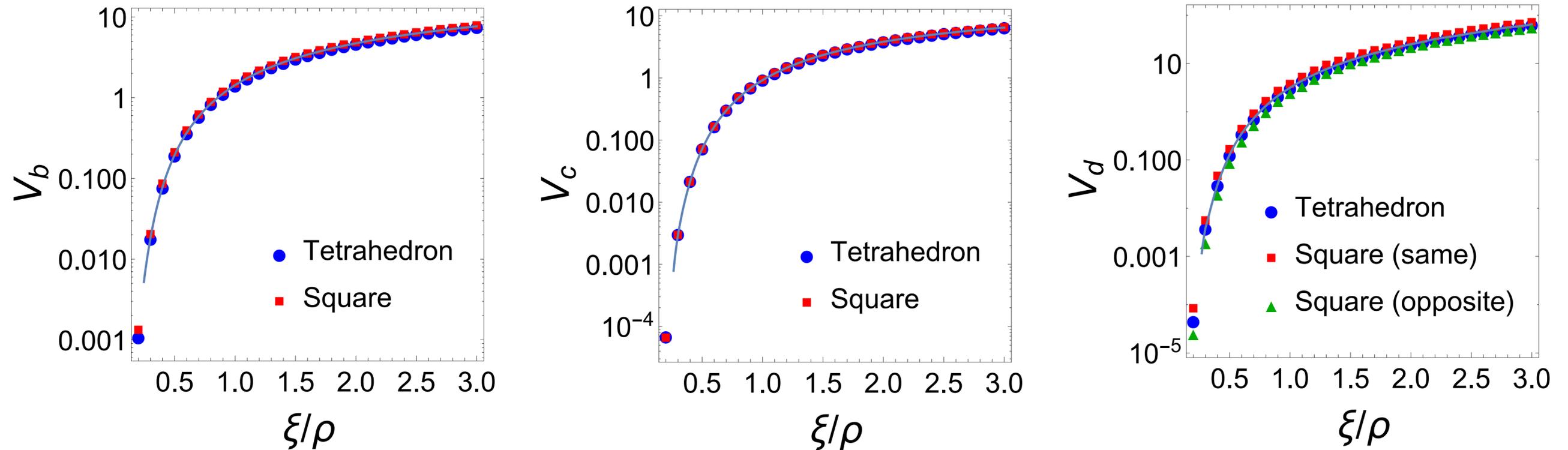


FIG. 3. (Color online) Interactions V_b (left), V_c (center), and V_d (right) corresponding to diagrams (b,c,d) of Fig. 2, respectively, as a function of the correlation-length-to-hyperdistance ratio ξ/ρ for both the tetrahedral and square configurations. The curve is an interpolation of the tetrahedral data points. The distinction between the 'same' and 'opposite' square configurations for diagram (d) is explained in the text.

At small correlation length very strong dependence on ξ
But it is moderate at $\xi/\rho > 1$

Rather weak dependence
on shape if ρ is the same

Averaging diagrams over snapshots from PIMC simulation

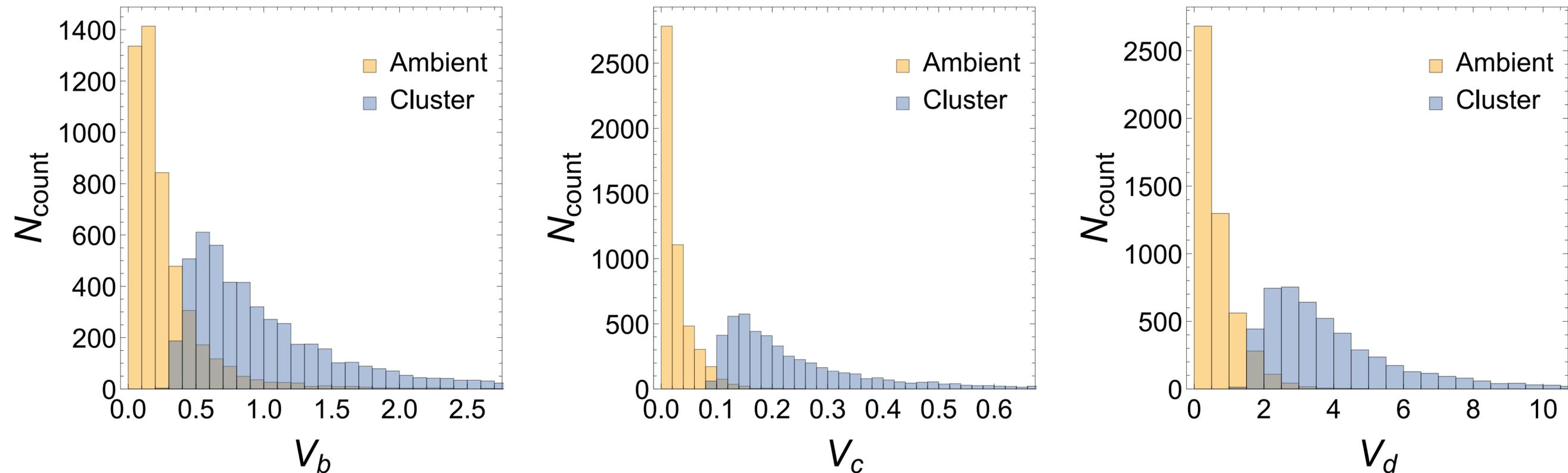


FIG. 6. Distribution of values of the multibody interactions V_b (left), V_c (center), and V_d (right) corresponding to diagrams (b,c,d) of Fig. 2, respectively, in 5000 configurations each for the cluster ($\rho < 3$ fm) and ambient nucleon matter ($\rho > 3$ fm) generated in PIMC simulation. Calculation performed with $\xi = 2$ fm.

Diagrams are significantly larger for clusters

Tails to the right are due to very small clusters: but those will be killed

Preliminary estimate: Landau ϕ^4 model

$$V_{tet} = -6 \frac{g_c^2}{4\pi} \langle V_a \rangle_{tet} + 4! \lambda_4 \left(\frac{g_c}{4\pi} \right)^4 \langle V_c \rangle_{tet}$$
$$\frac{g_\sigma^2}{4\pi} = 6.04, \quad \frac{g_\omega^2}{4\pi} = 15.17.$$

Walecka model of
Relativistic mean field
For nuclear matter

Critical mode is their mixture
Stephanov used 10
as some round average
And so do we

But we do not know
The value of
The quartic coupling

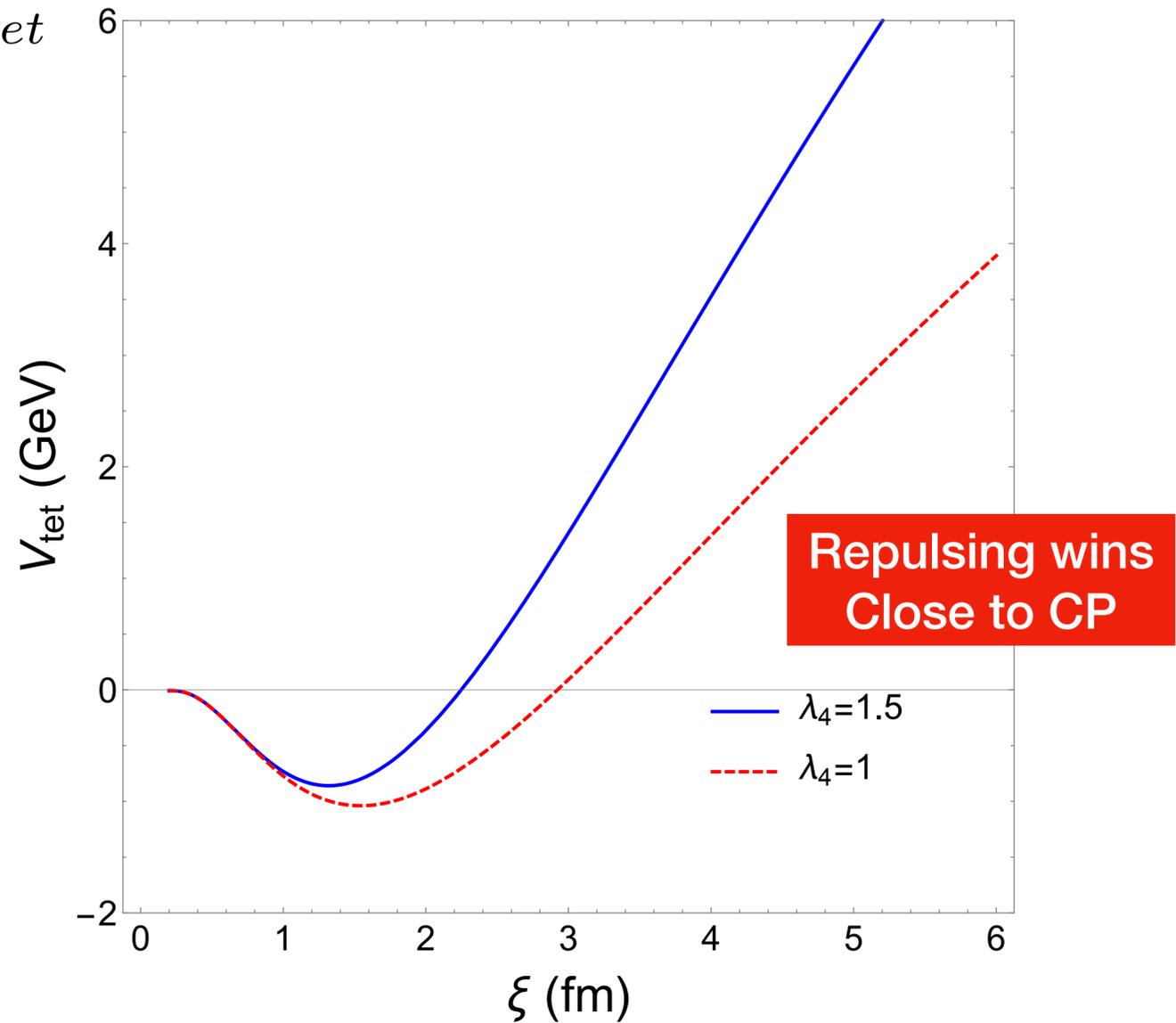


FIG. 4. Energy of four-nucleon tetrahedral cluster (in GeV) as a function of correlation length ξ (fm). The critical mode-nucleon coupling is taken to be equal to nucleon-sigma meson coupling of the Walecka model (20), and the values of quartic coupling $\lambda_4 = 1.5$ (upper curve) and $\lambda_4 = 1$ (lower curve).

THE UNIVERSAL EFFECTIVE ACTION FOR ISING-TYPE CRITICAL FLUCTUATIONS

The Landau model, used as an initial approximation, does *not* however represent correct behavior near Ising-like critical points. Wilson's epsilon expansion – in $\epsilon = 4 - d$ where d is space dimension – has found that under the renormalization group flow the Landau model goes into the fixed-point regime in infrared, with small coupling at small ϵ . While Wilson famously calculated approximate values of the critical indices, one might still doubt whether ϵ -expansion gives an accurate account at $\epsilon = 1, d = 3$.

Three arguments suggesting that at the critical point $\Omega \sim \phi^6$.

1. $\frac{d\Omega}{d\phi} = J. \quad \langle \phi \rangle(J) \sim J^{1/\delta}, \quad \delta = \frac{d + 2 - \eta}{d - 2 + \eta} \approx 4.78 \quad \text{Much closer to 5 than to 3}$

2. including a ϕ^6 term – **but not higher powers** – can be justified because this term is the renormalizable one, in $d = 3$ case

3. there are numerical studies showing this ansatz for Ω gives good fits of lattice data.

Probing effective action at large phi is done by
 doing simulations with different J
 At critical line parameterized by $t=T/T_c-1$

$$\Omega(\phi) = \int d^3x \left[\frac{(\phi_{,\mu})^2}{2} + \frac{m^2\phi^2}{2} + mg_4\phi^4 + g_6\phi^6 \right]$$

Note that at $m \rightarrow 0, \xi \rightarrow \infty$ it indeed has only the last ϕ^6 term.

Dimensionless
 couplings are fixed

$$m = \frac{1}{\xi} \approx Mt^\nu \quad \tilde{\phi} \equiv \frac{\phi}{M}$$

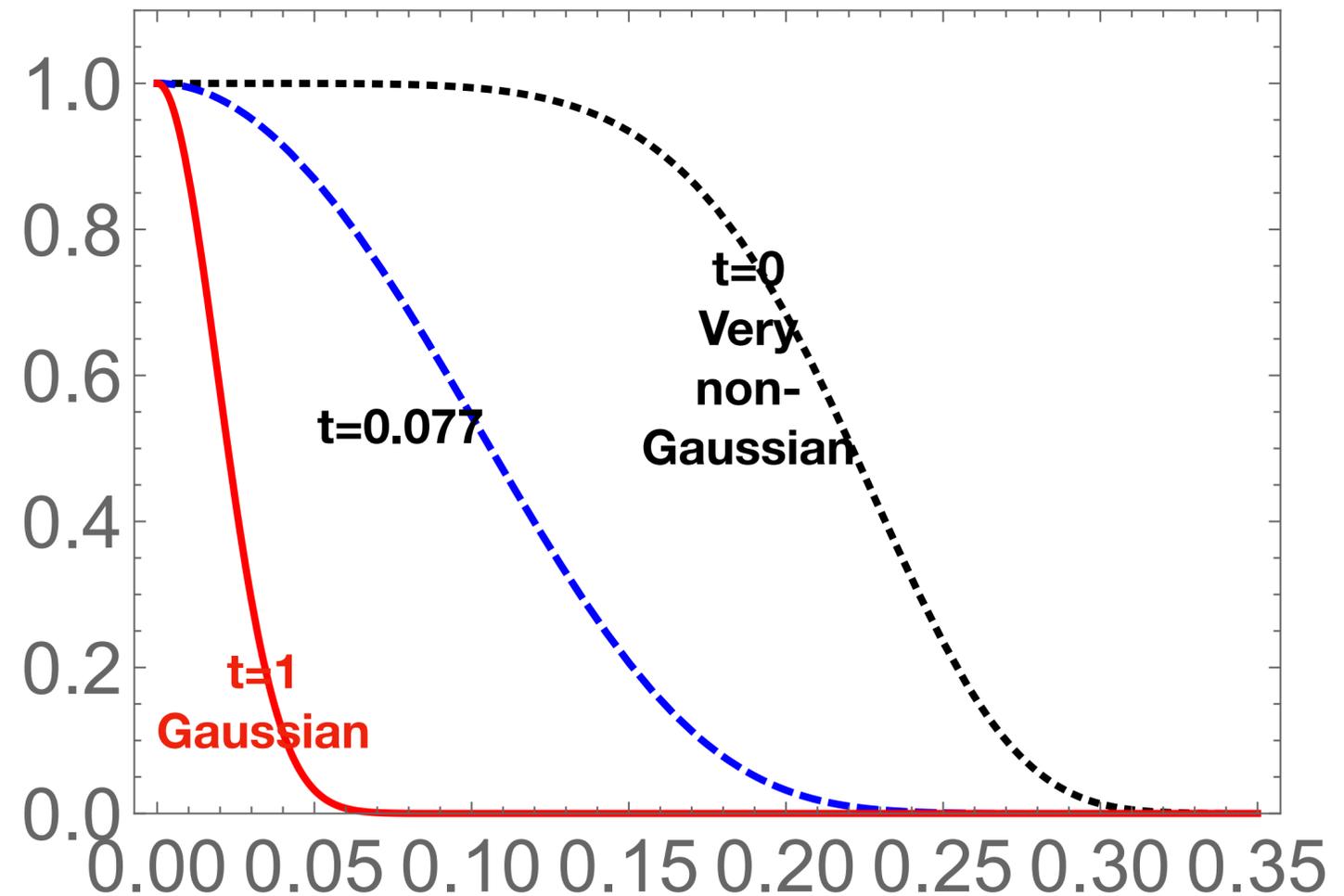
at CP, $\xi = \infty, t = 0$, only the last term survives.
 Only one dimensional parameter M
 Which we take to be sigma mass

$$Z = \int D\phi e^{-(\Omega(\phi) + J(x)\phi(x))V_3/T}$$

g4=0.97, g6=2.05 from fit

M. M. Tsy-pin, (1994), arXiv:hep-lat/9401034 [hep-lat].
 Agrees also with RG calculation by Heidelberg group

P=exp(-VM^3 Omega)



DEFORMED EFFECTIVE POTENTIAL NEAR THE CRITICAL LINE

$$\frac{\partial \Omega}{\partial \tilde{\phi}}(\tilde{\phi}_0) = J$$

As example we use dimensionless $J = 1/100$
 Then solve the 5-th order eqn for maximum
 Then re-center the distribution by

$$\tilde{\phi}_0(t = 0.01) \approx 0.224, \quad \tilde{\phi}_0(t = 0.41) \approx 0.031$$

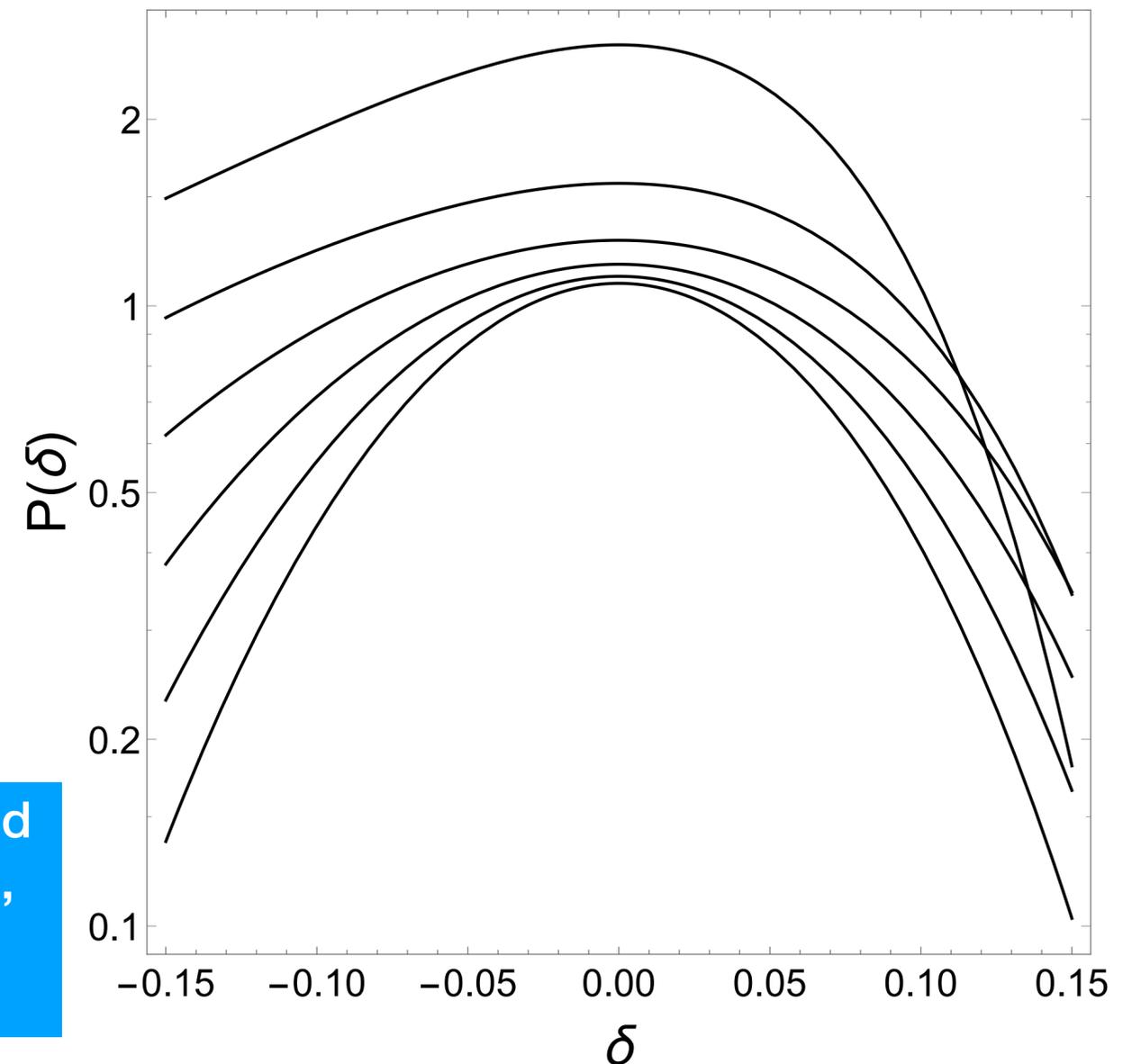
And get new action in terms of delta
 Which has all powers of it except the first

$$\tilde{\phi} = \tilde{\phi}_0 + \delta$$

$$\Omega_{def}(t = 0.01) \approx -0.0017 + 0.095\delta^2 + 0.51\delta^3 + 1.60\delta^4 + 2.75\delta^5 + 2.05\delta^6$$

No linear term, small quadratic one => xi not infinite even at t=0

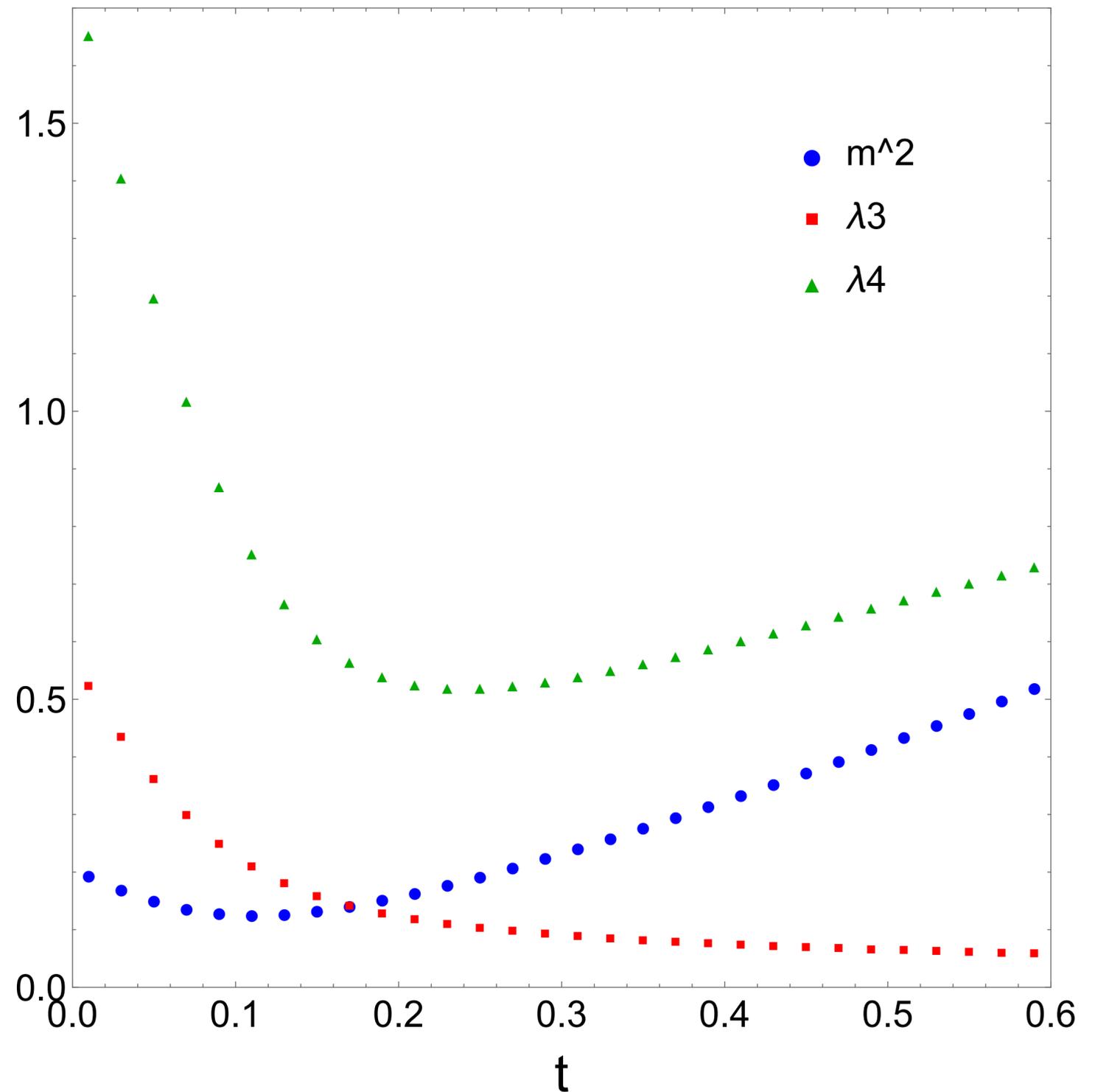
Six curves, top to bottom, correspond to values of $t = 0.01, 0.09, 0.19, 0.29, 0.39, 0.49, 0.59$.



DEFORMED EFFECTIVE POTENTIAL NEAR THE CRITICAL LINE

Triple and
Quartic
couplings
Strongly
grow
Near CP, $t \rightarrow 0$

$m=1/\xi$ does not vanish
Near CP but remains small

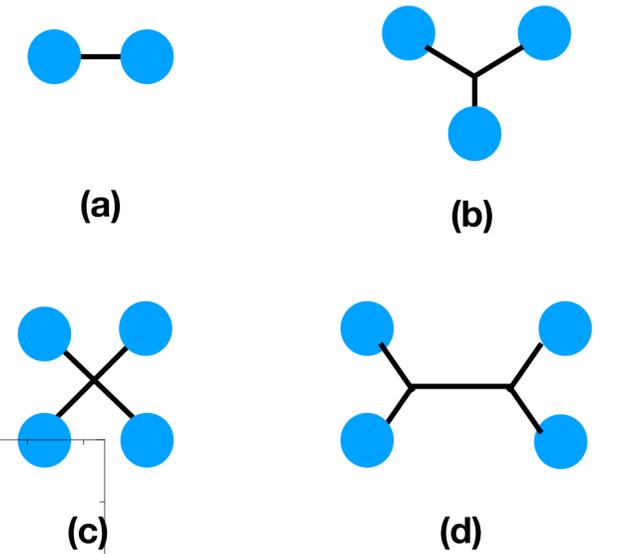


Taking all effects together

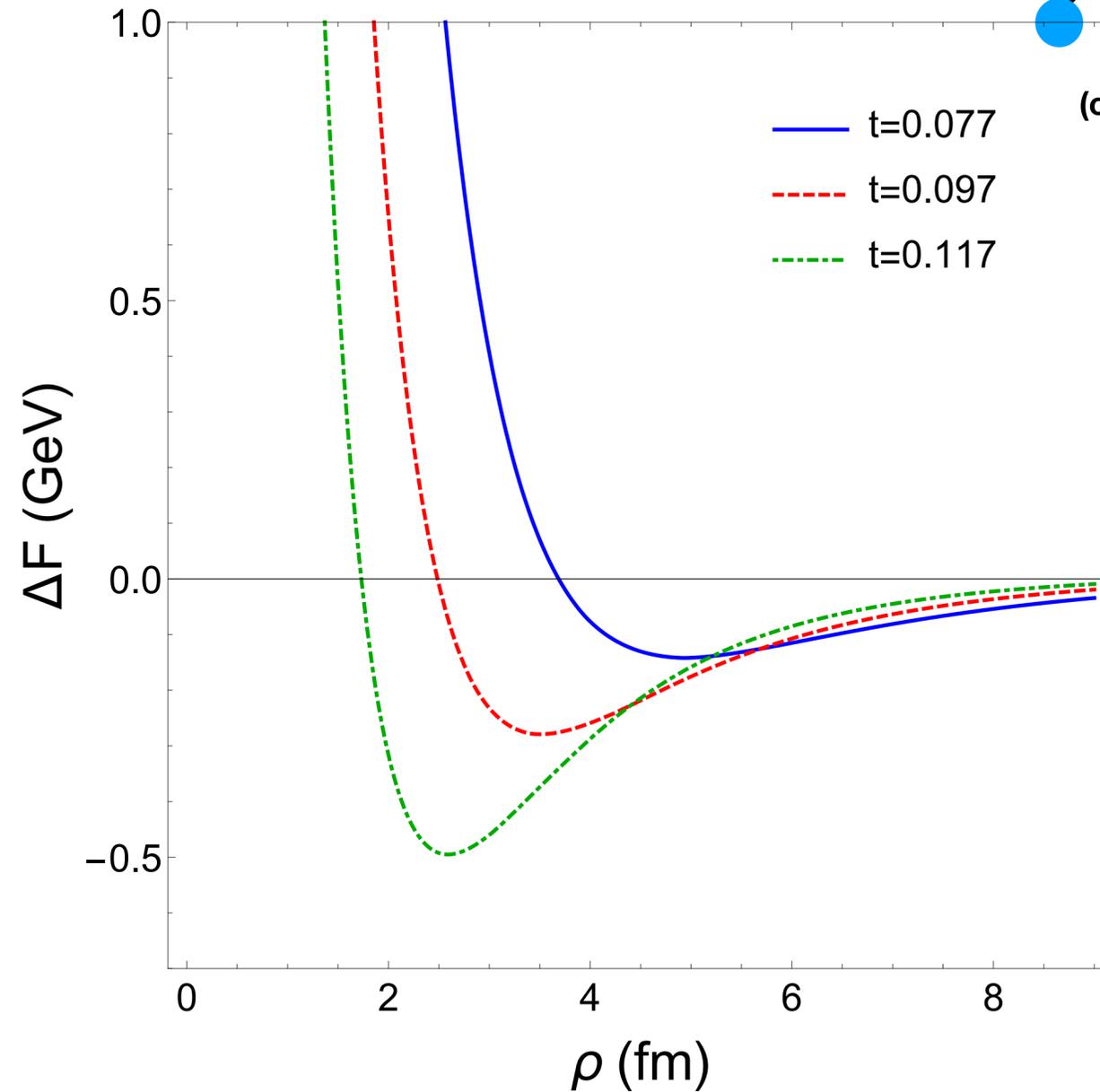
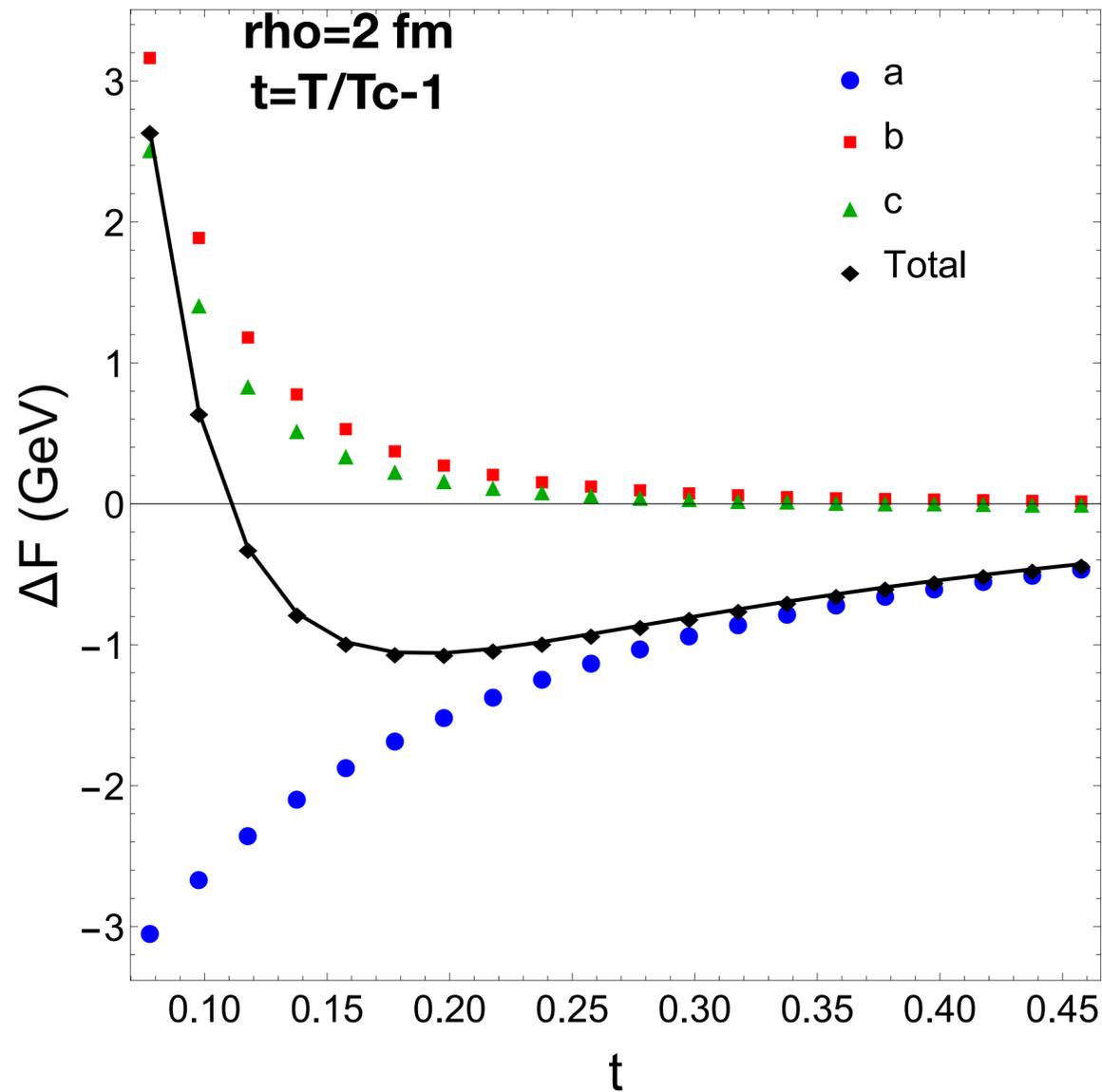
Repulsive three and four
Body forces overcome
Attraction near CP

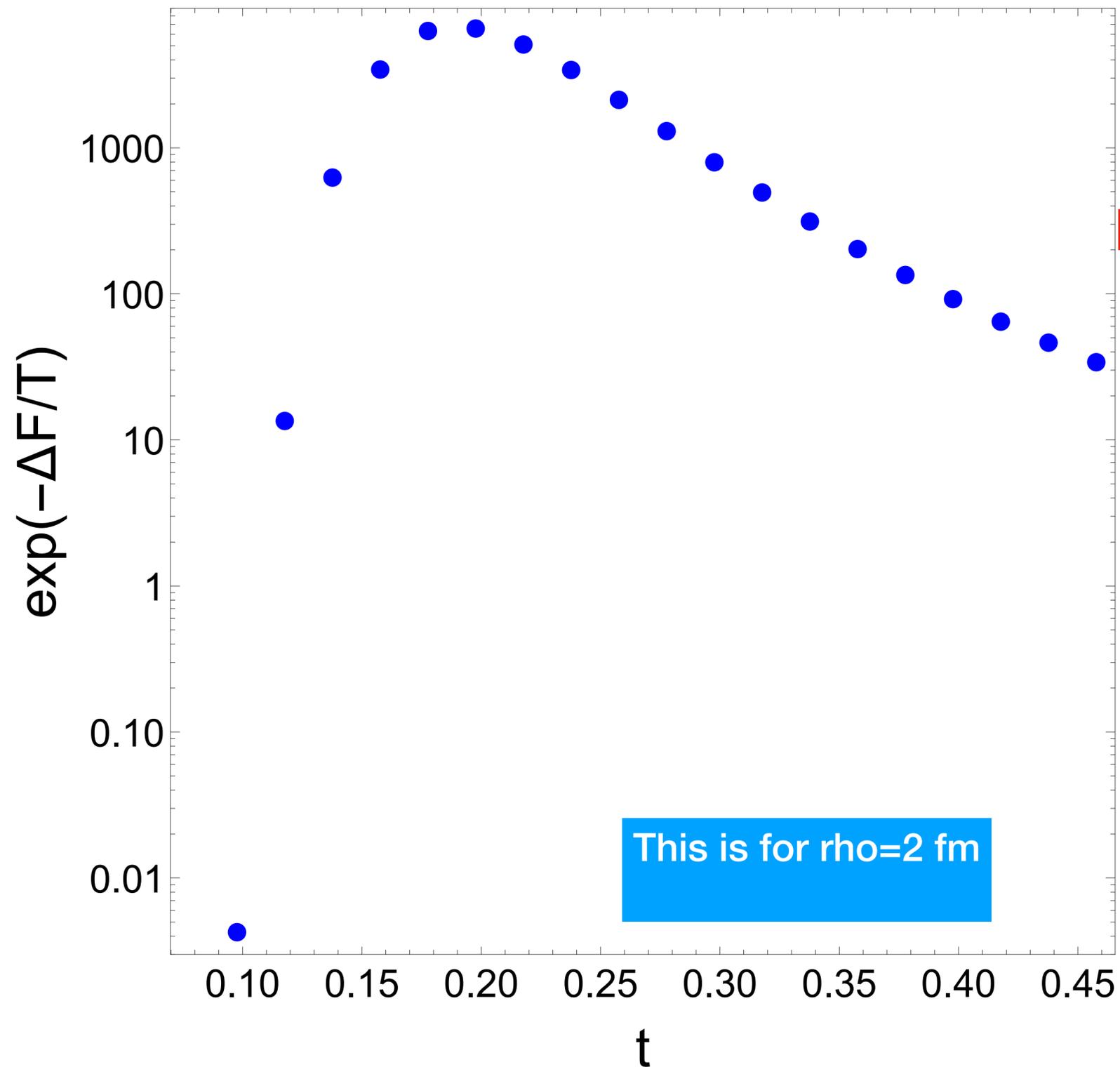
$$V = -\frac{4 \cdot 3}{2} \frac{g_c^2}{4\pi} \frac{\exp(-r_{ij}/\xi)}{r_{ij}} + 4 \cdot 3! \lambda_3 \left(\frac{g_c}{4\pi}\right)^3 V_b$$

$$+ 4! \lambda_4 \left(\frac{g_c}{4\pi}\right)^4 V_c - 4! \frac{\lambda_3^2}{8\pi} \left(\frac{g_c}{4\pi}\right)^4 V_d$$



Energy of 4-N cluster





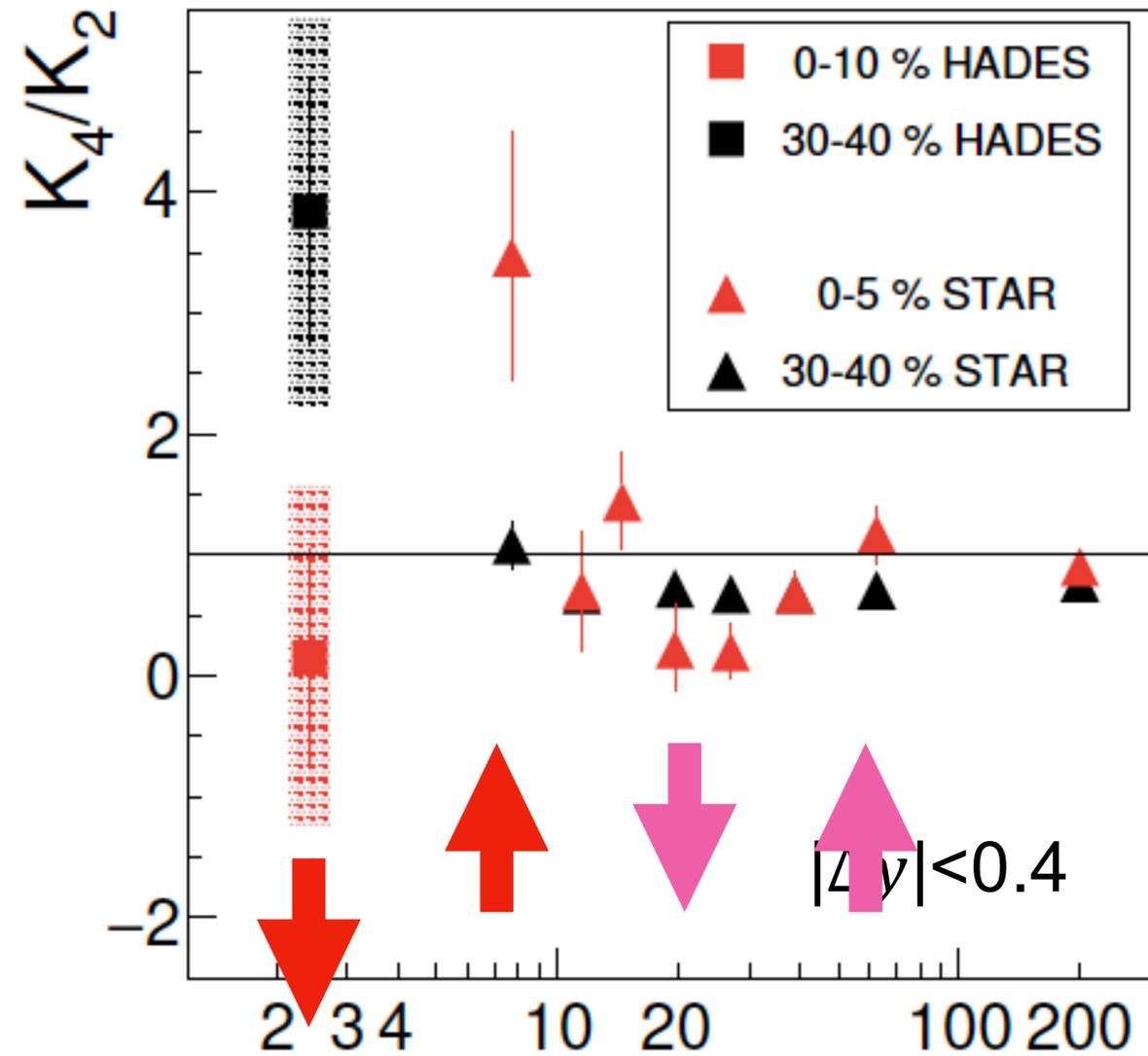
Plotting $\exp(-V/T)$
One gets dramatic
non-monotonous signal
for cluster formation

Perhaps the nucleon
coupling to critical mode
is not that large as assumed
But qualitative shape
is now clear

One needs to look for a dip
in clustering

Let us now look at experimental kurtosis

Older STAR data have shown large effect



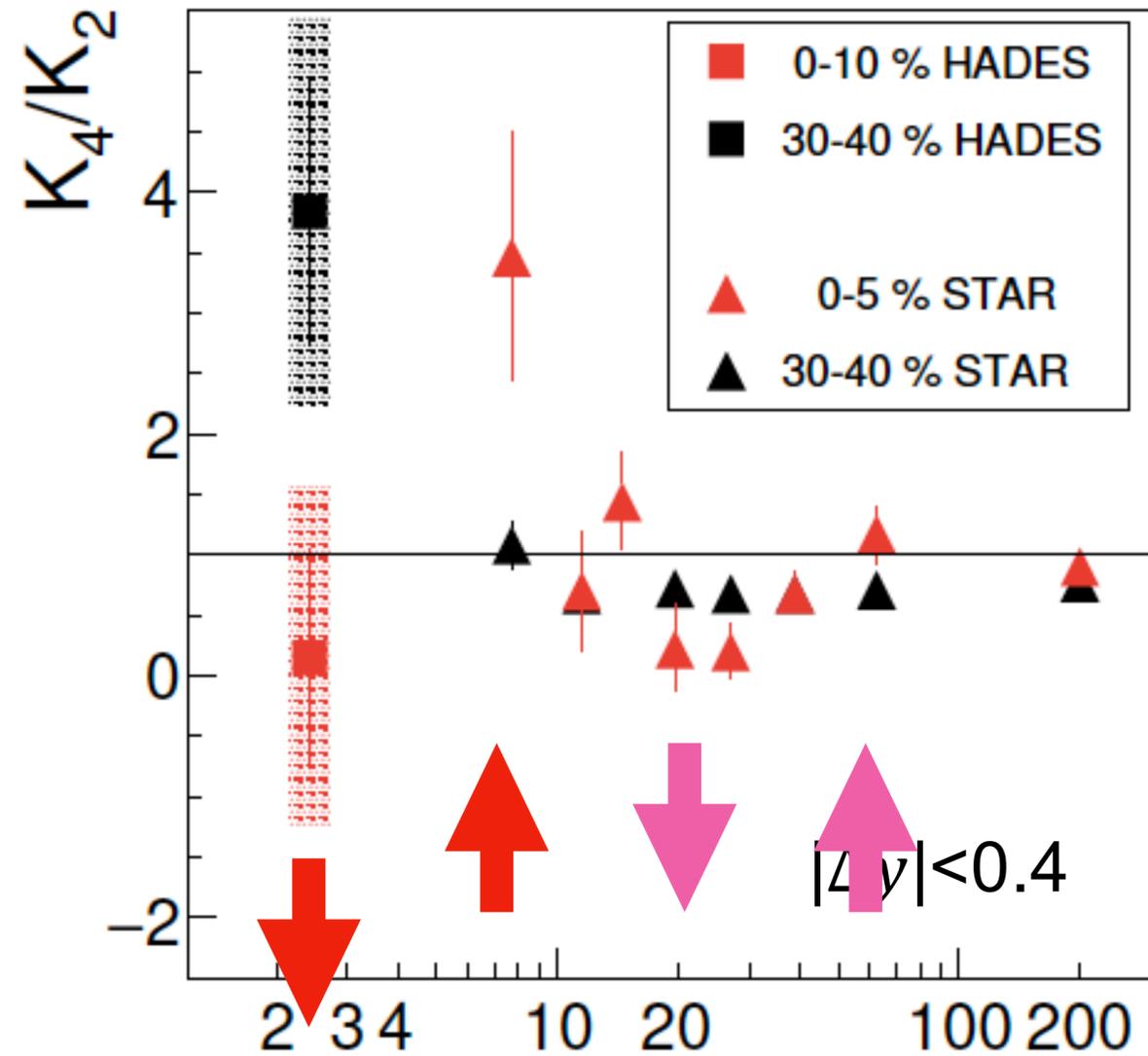
• e-Print: [2001.02852](#)

Two dips for central bins
large at 2 and smaller at 20 GeV?
Errors still large => BESII

Let us now look at experimental kurtosis

Older STAR data have shown large effect

Which was recently found to be partly due to small set of defective events for central bin

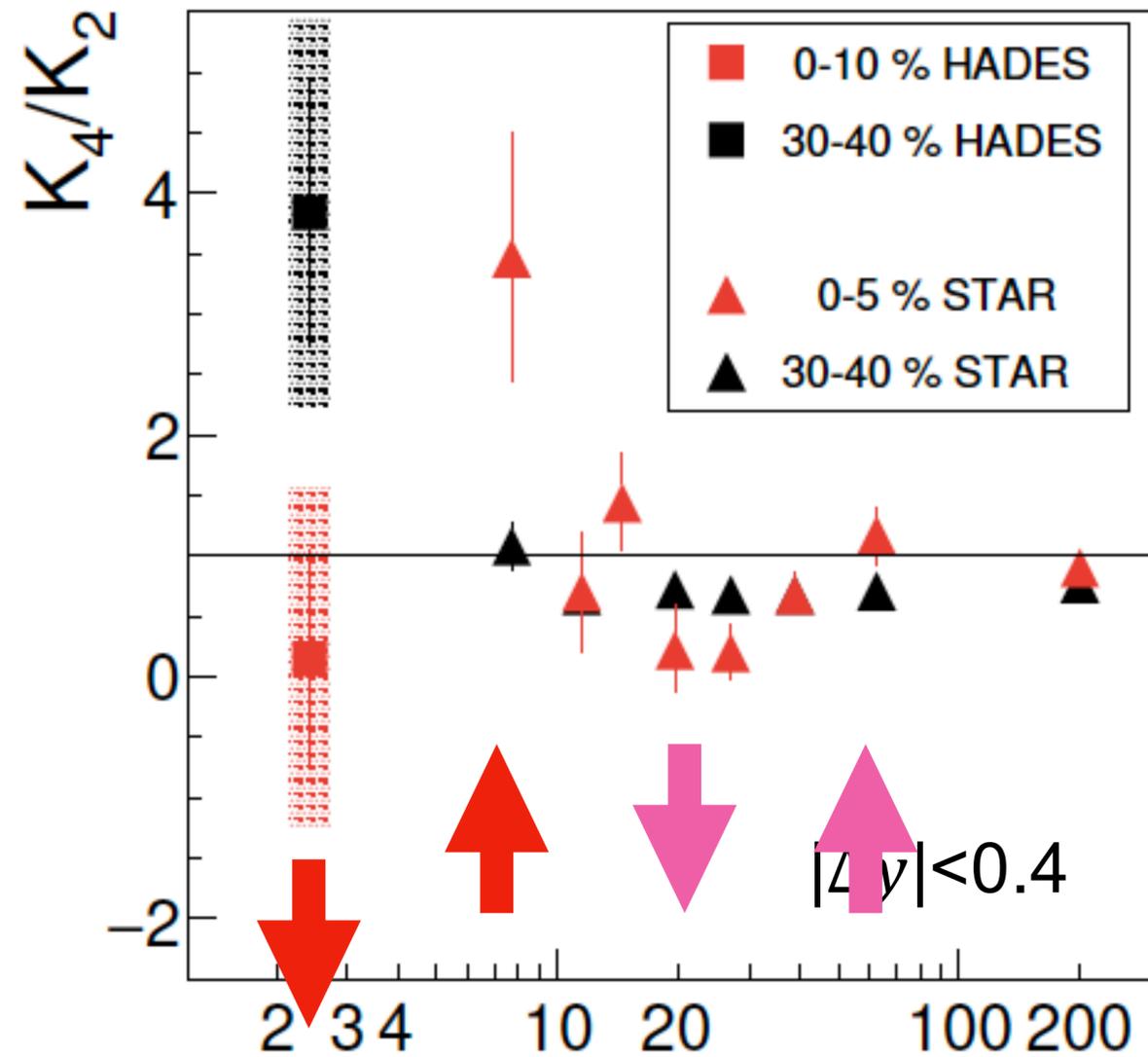


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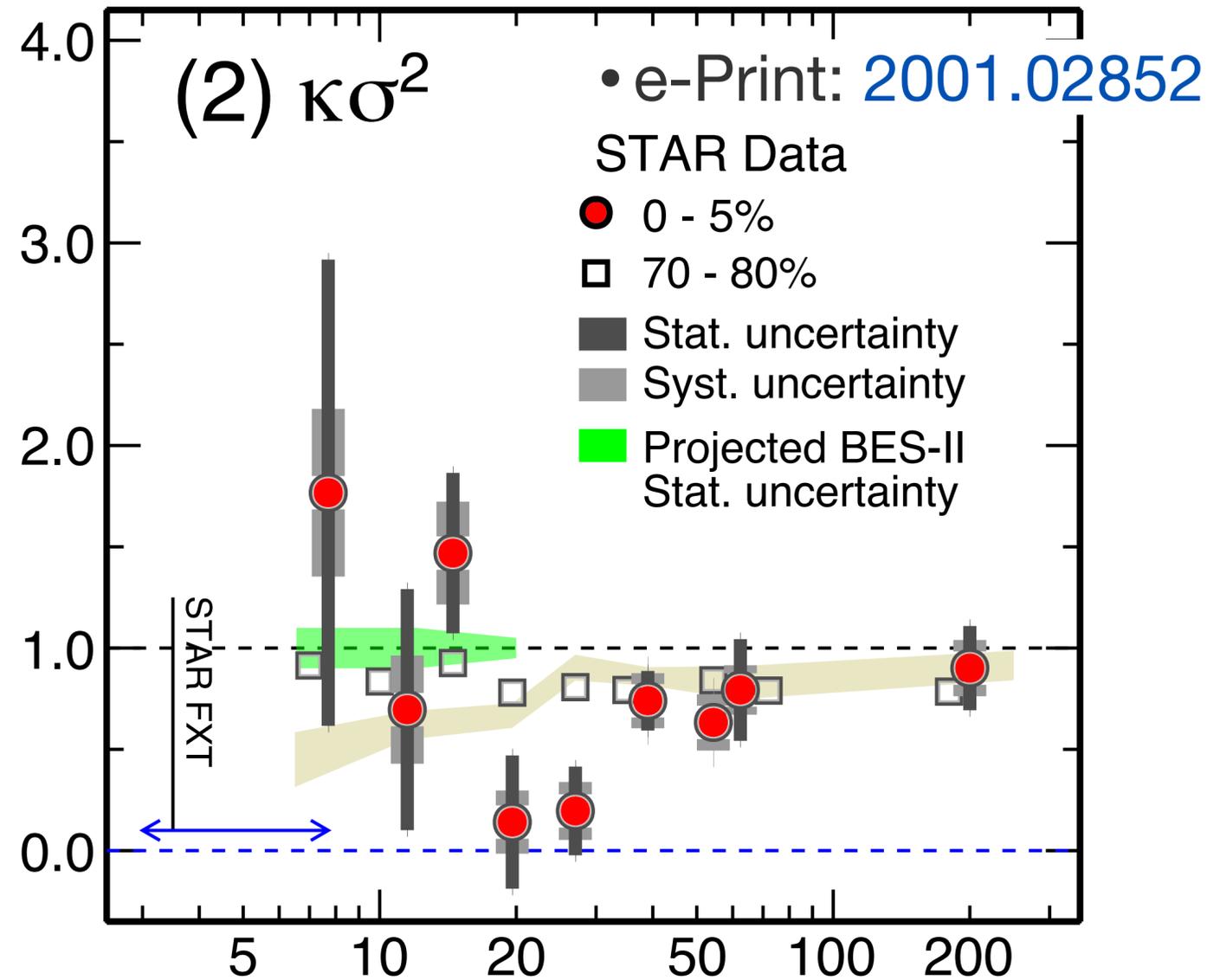
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(2) $\kappa\sigma^2$

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STAR Data

● 0 - 5%

□ 70 - 80%

■ Stat. uncertainty

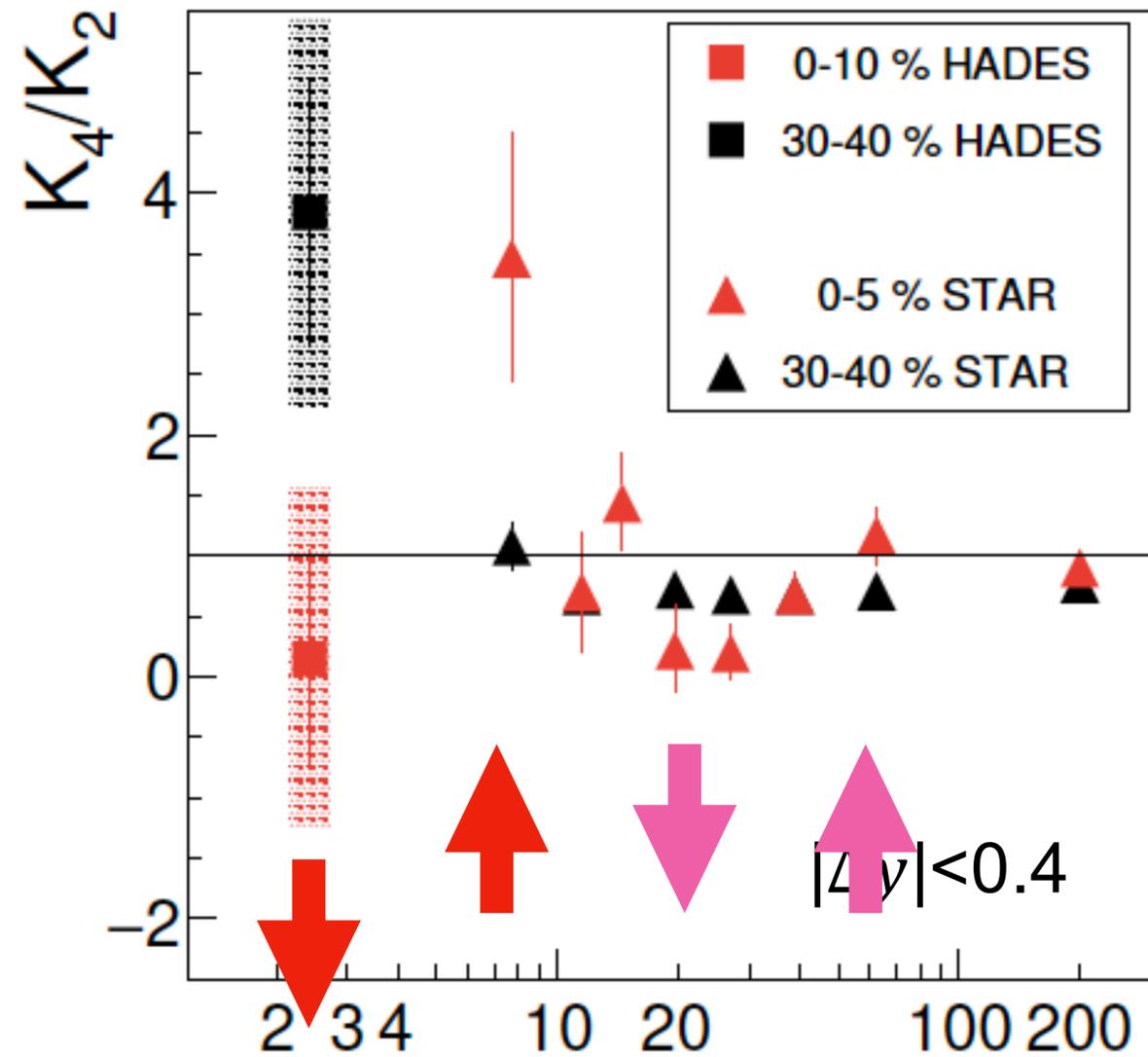
■ Syst. uncertainty

■ Projected BES-II
 Stat. uncertainty

STAR FXT

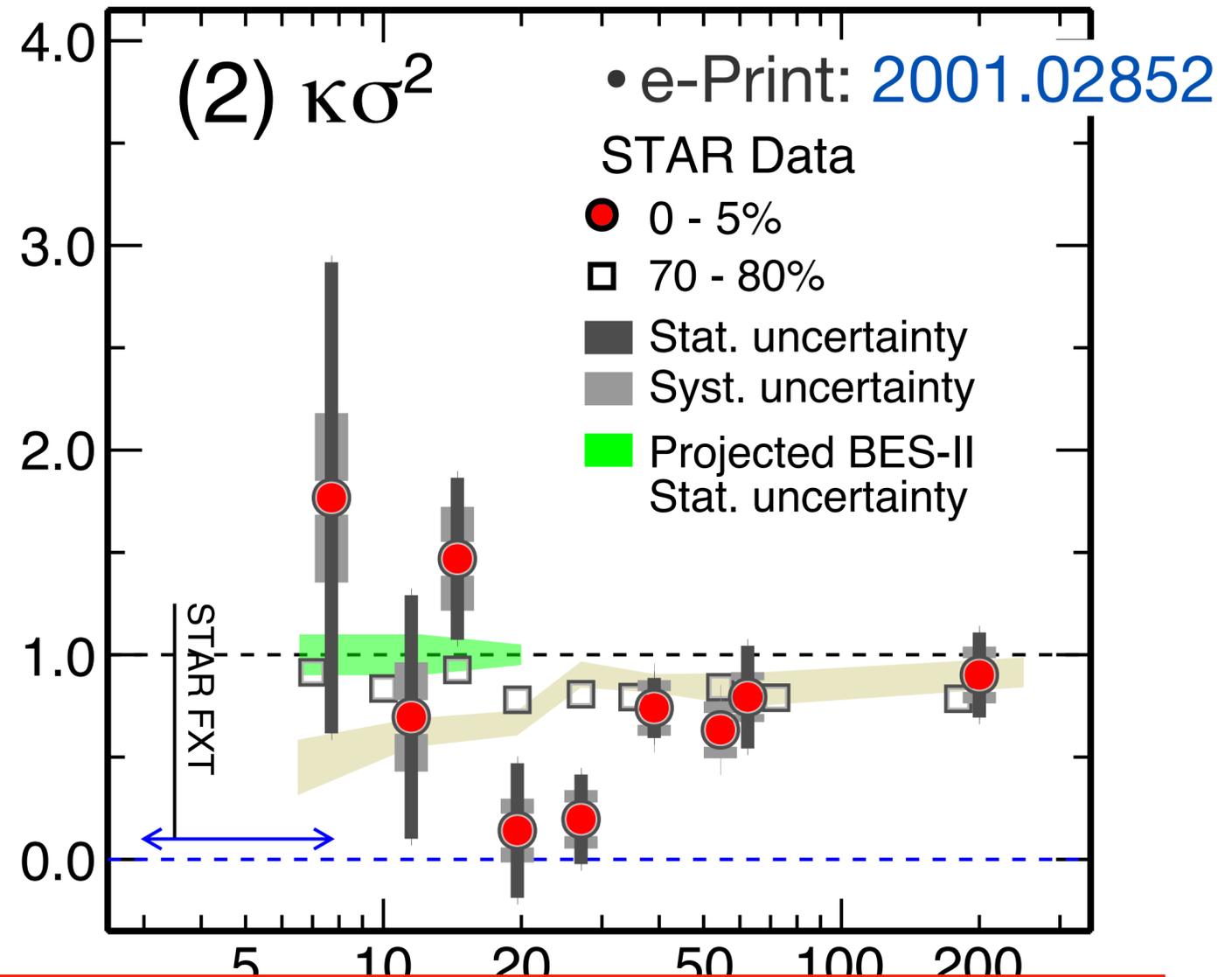
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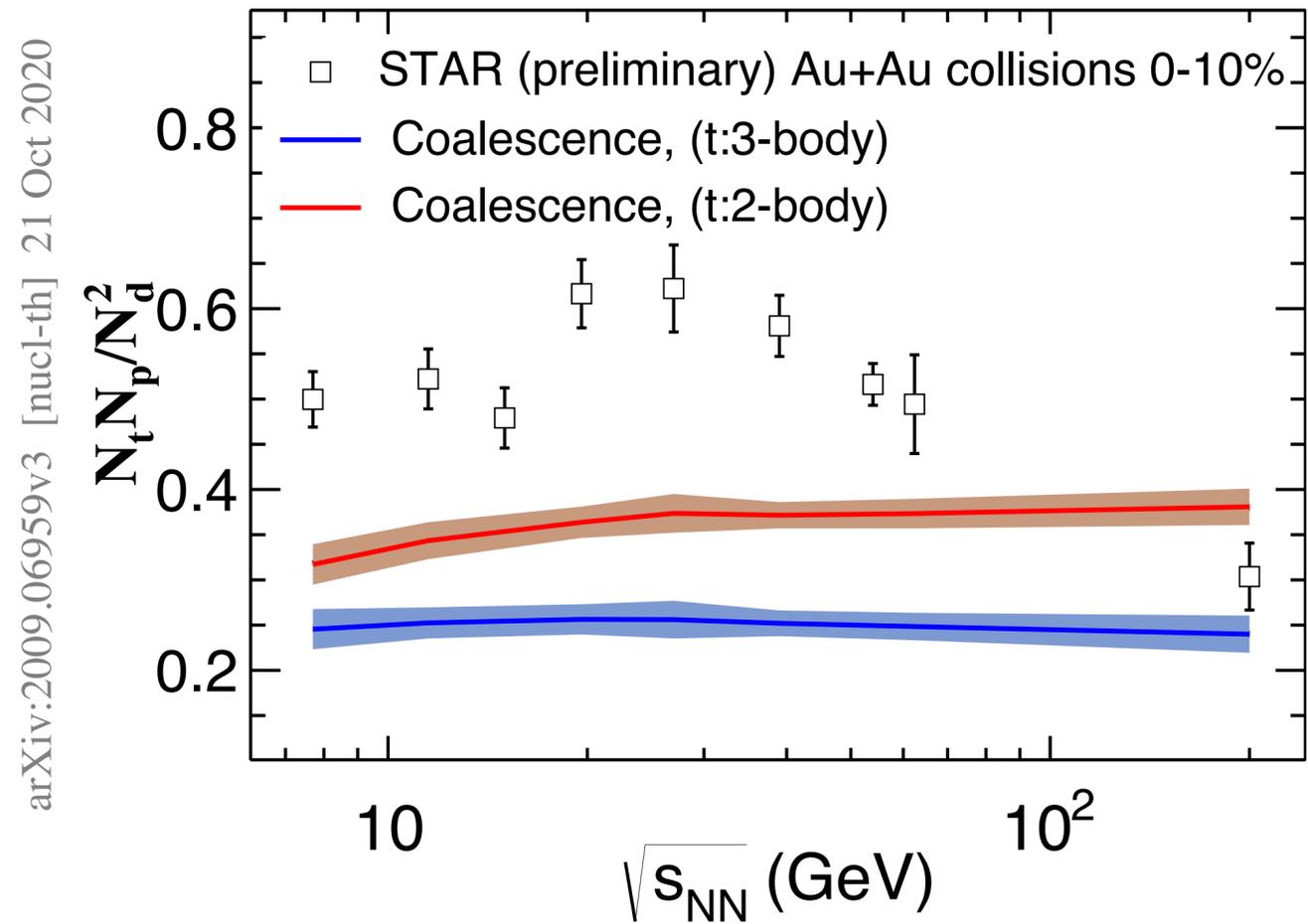
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Still non-monotonous signal?
Clearly much more accurate measurements
from BeS-II are needed

Let us now look at light nuclei production: the tritium ratio

In this ratio the main driver
— fugacity $\exp(\mu/T)$ —
Cancels out



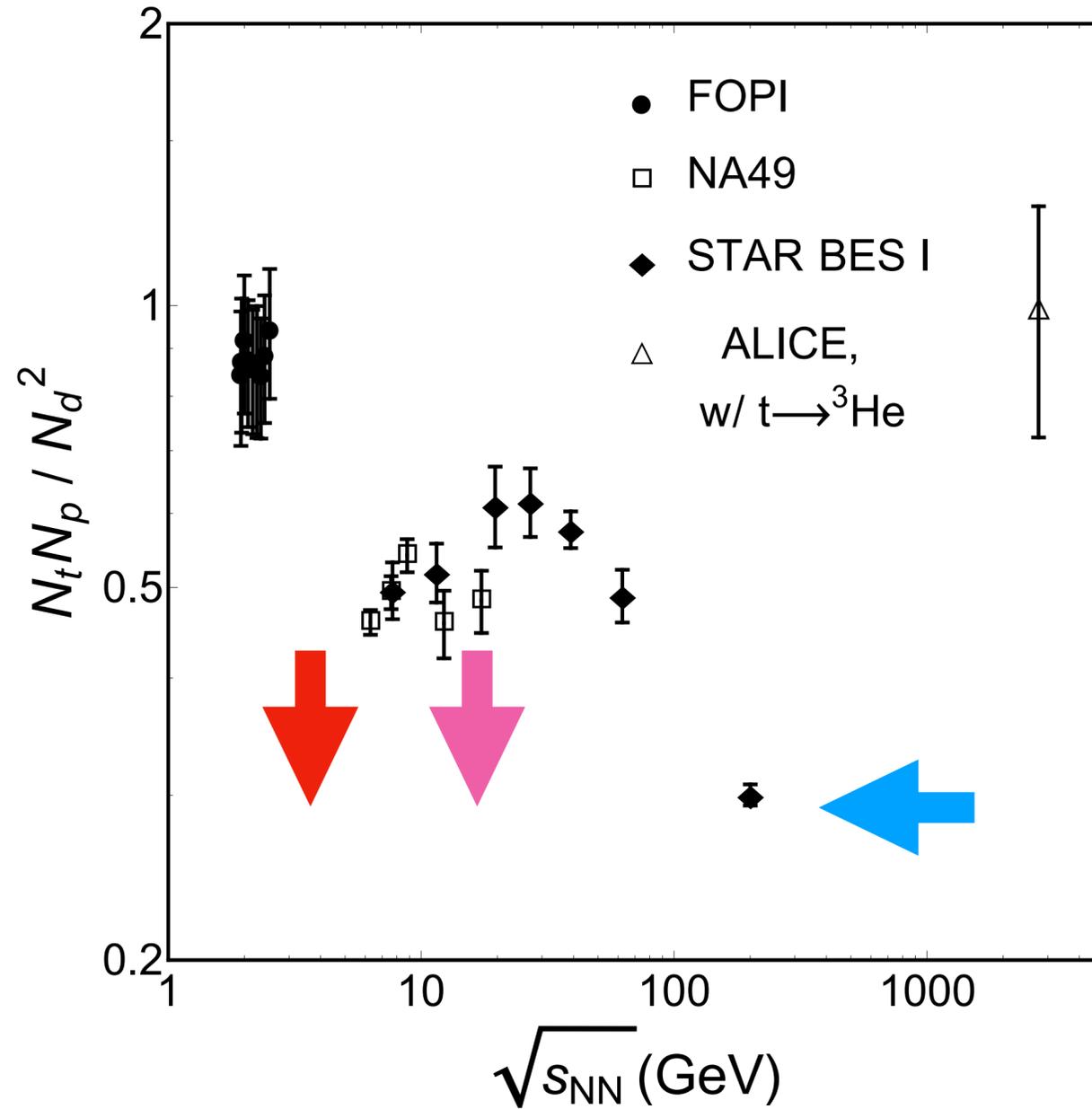
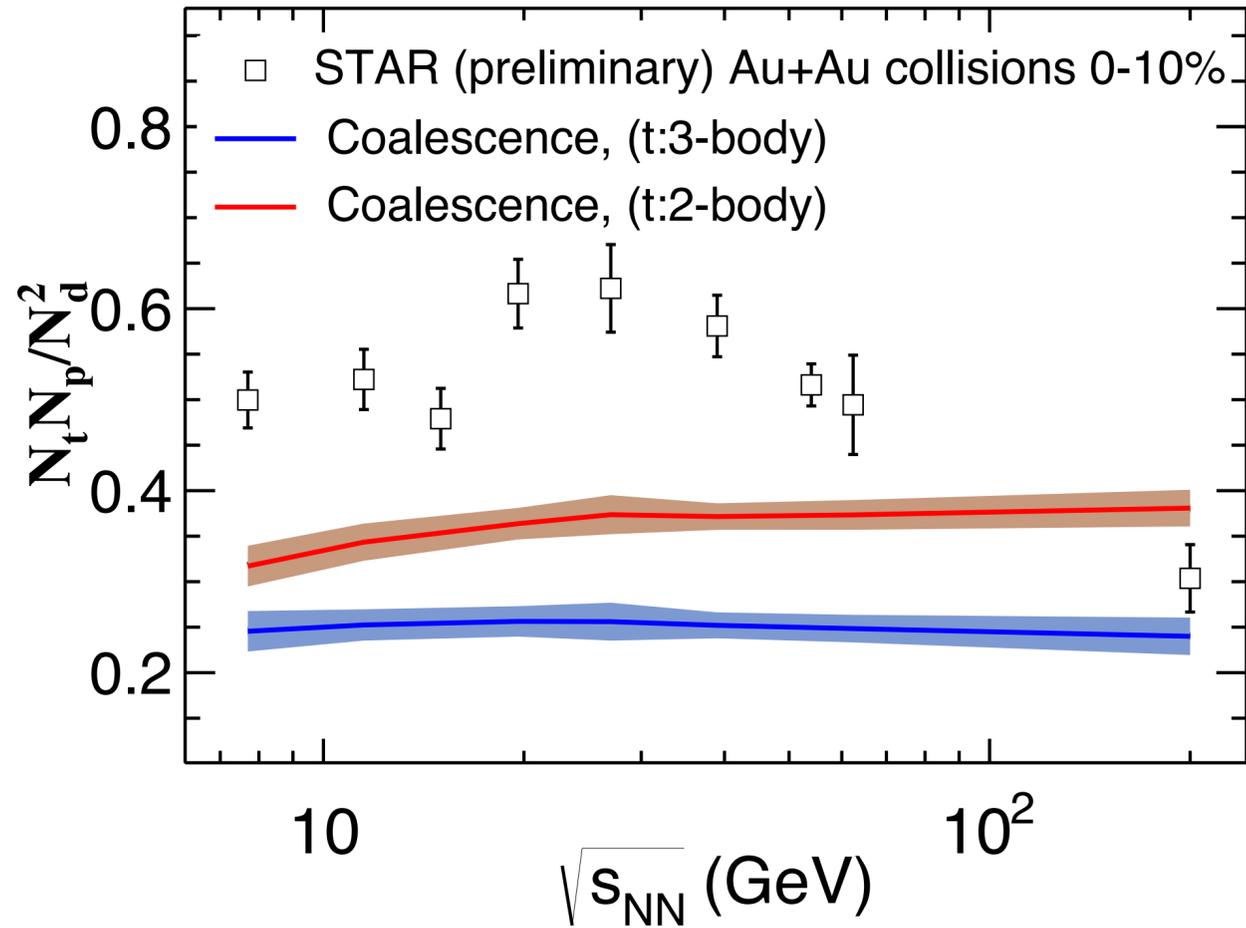
Wenbin Zhao,^{1,2,3,4} Chun Shen,^{5,6} Che Ming Ko,⁷ Quansheng Liu,^{1,2} and Huichao Song^{1,2,3}

**Extra source of t is needed:
4-N reclusters =>
50 states of He4 =>
feeding**

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arXiv:2009.06959v3 [nucl-th] 21 Oct 2020



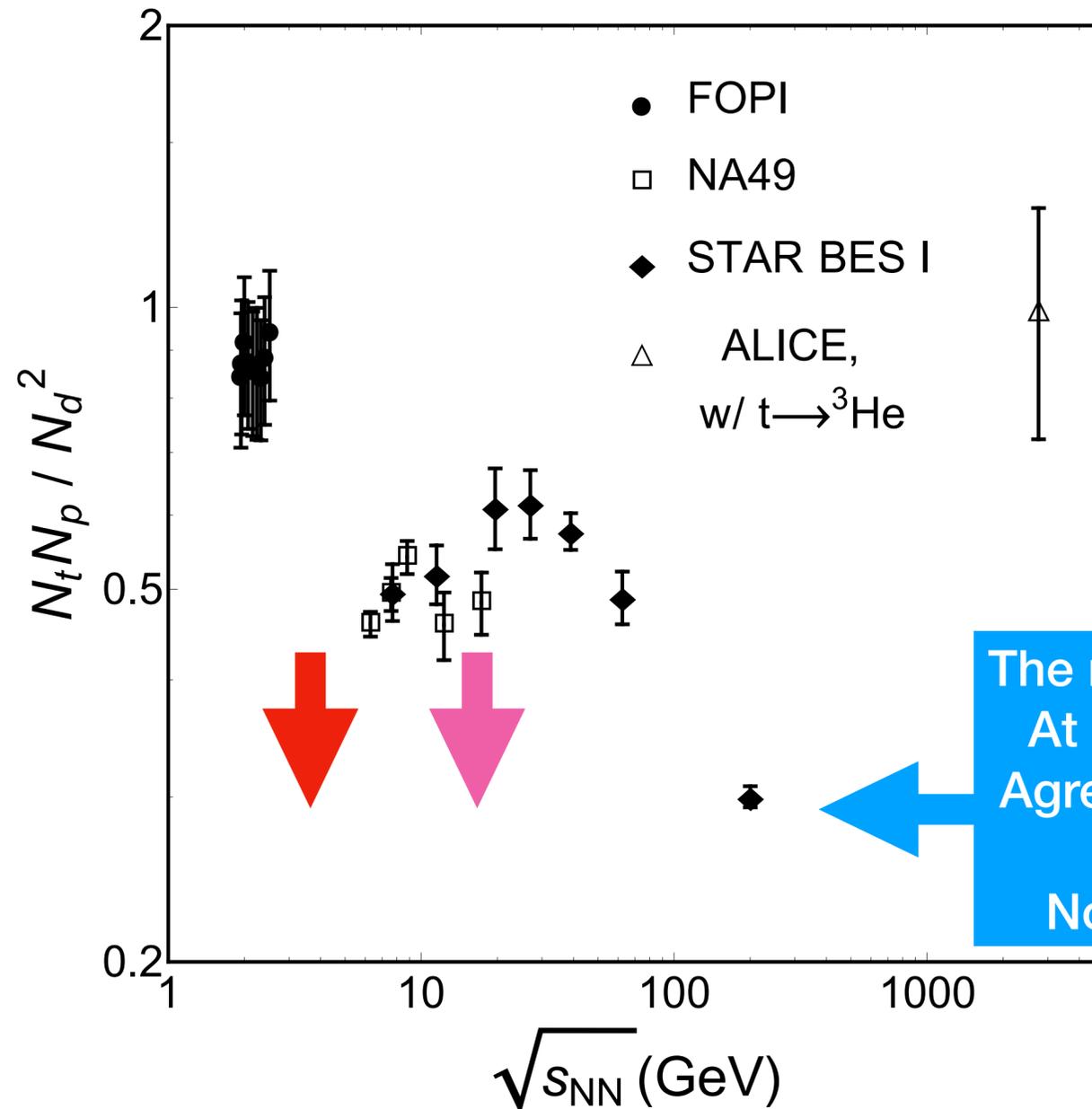
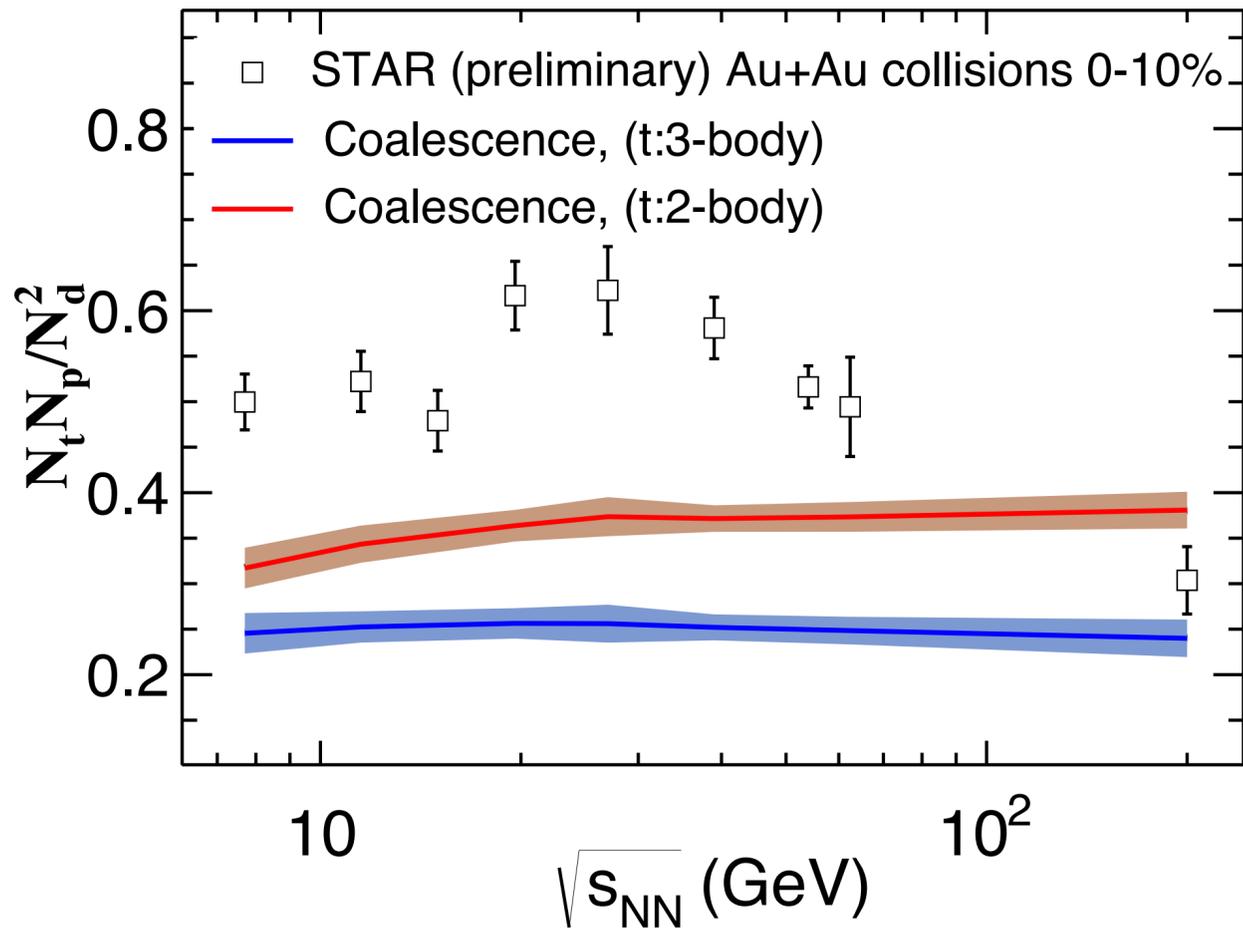
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The most accurate point
 At main RHIC energy
 Agrees with stat weight
 Ratio 0.29 =>
 No feeddowns here?

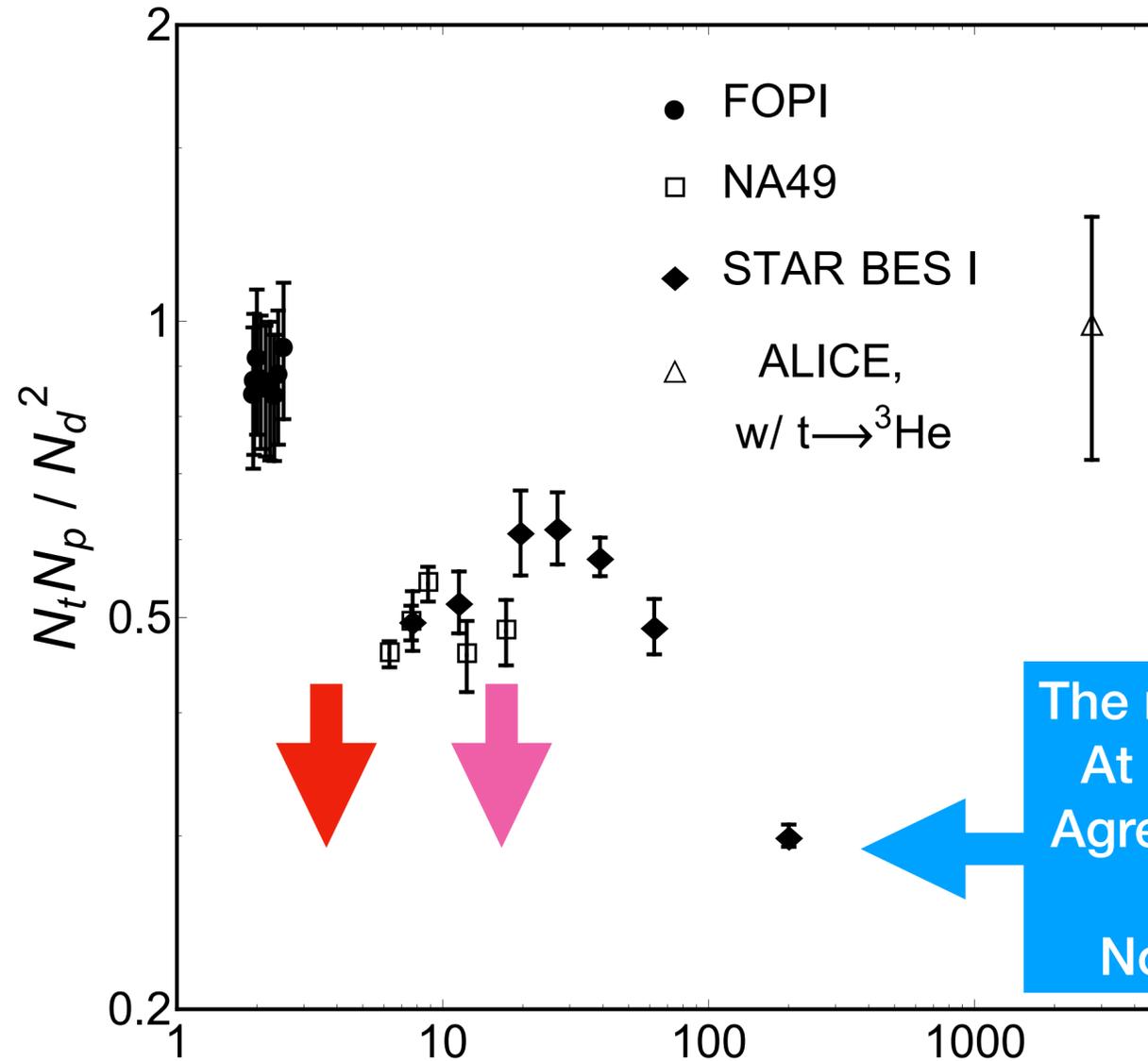
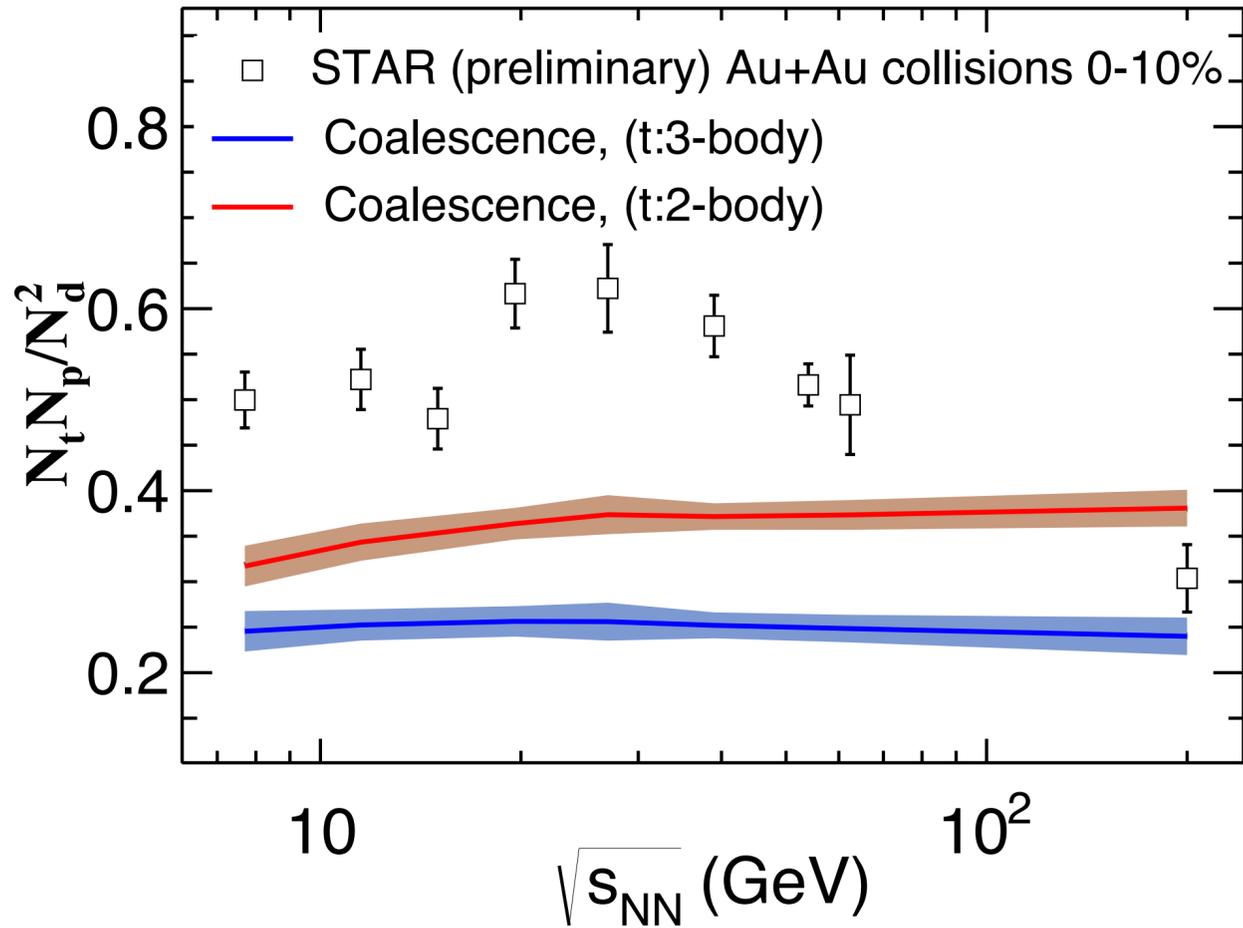
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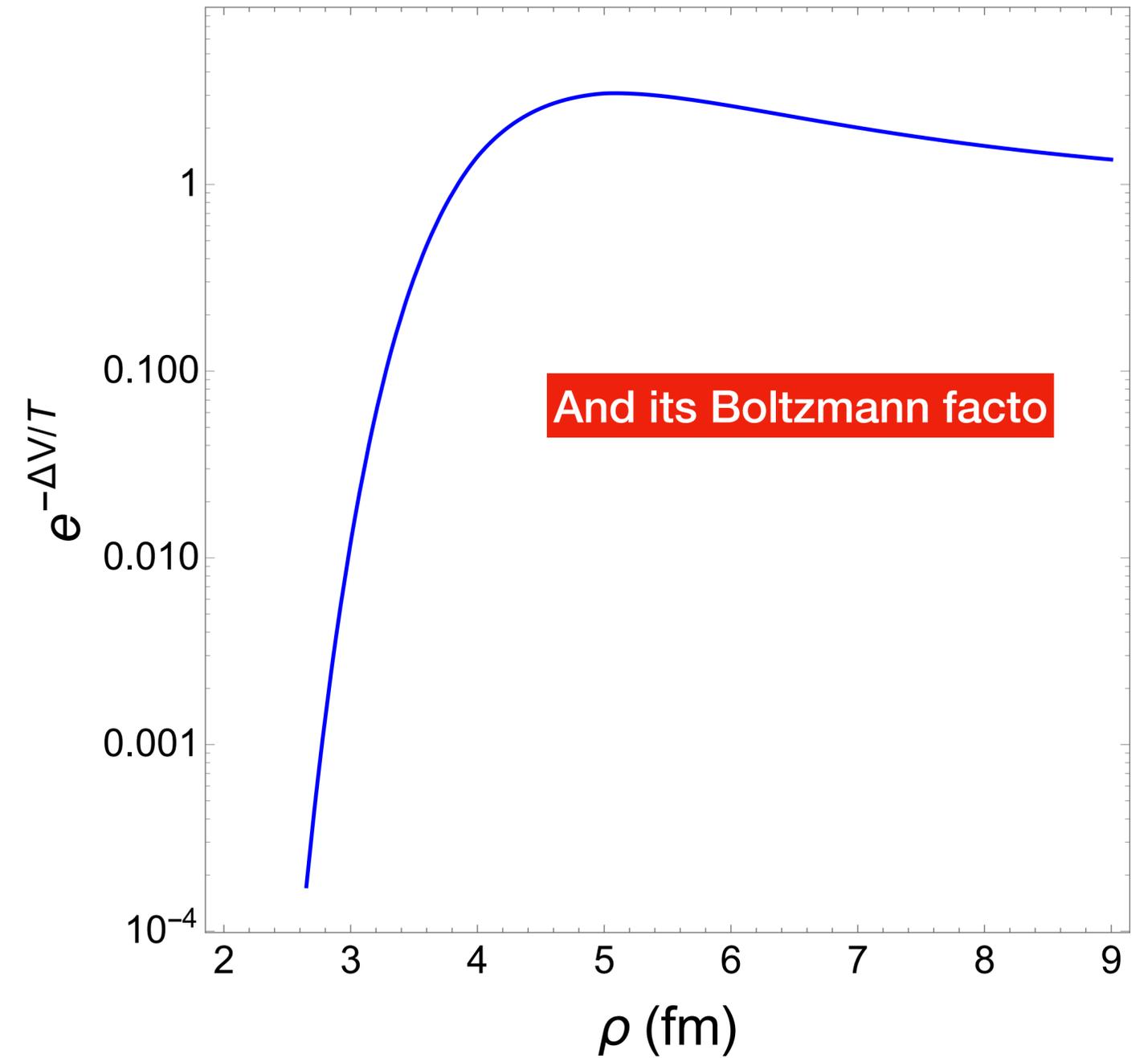
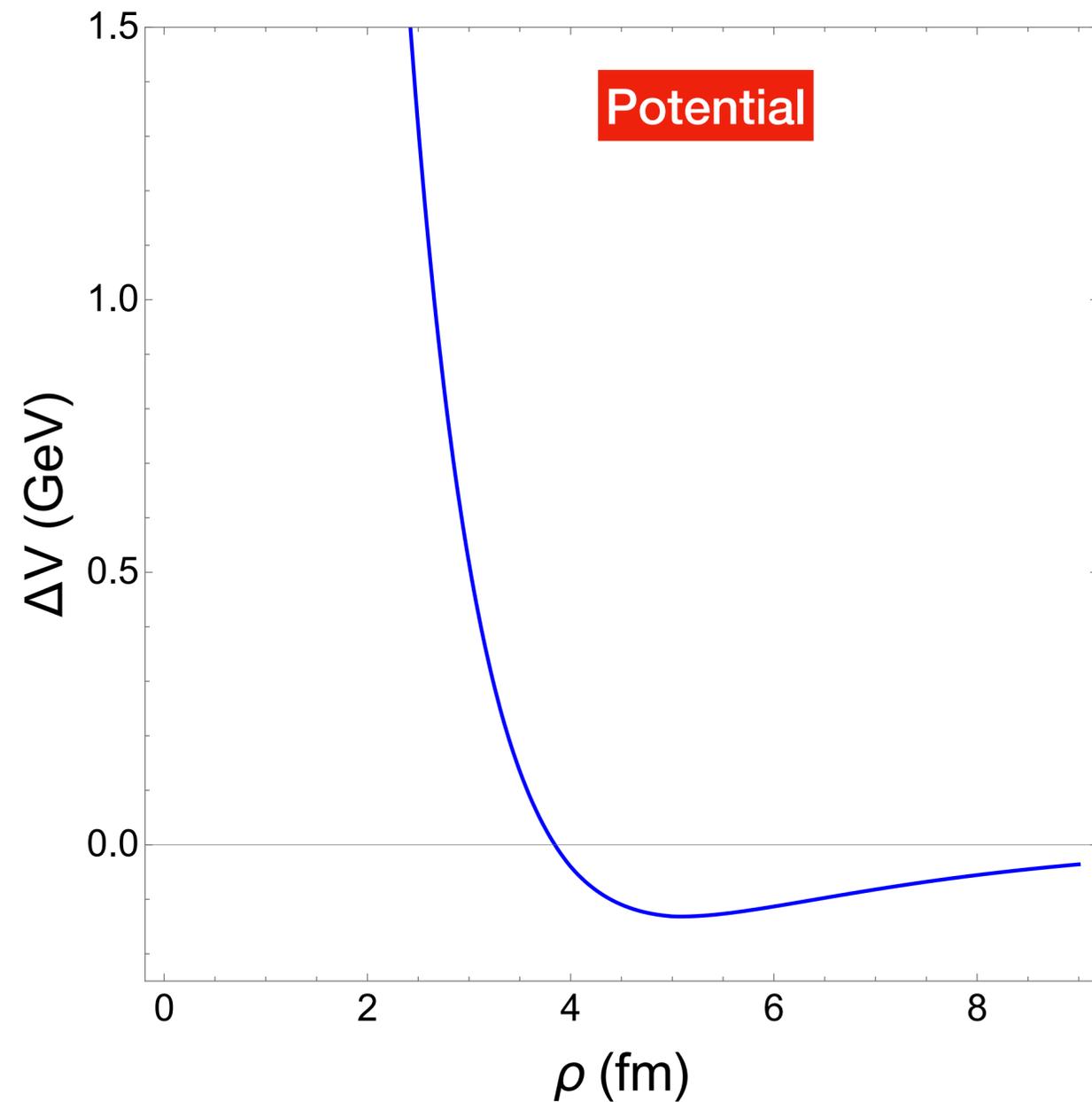
Also two dips for central bins
 large at 2 and smaller at 20 GeV?
 Errors still large => BESII

Summary

- **Paradox (ES,2006) at CP is not there** even at $\xi \rightarrow \infty$ there is no implosion
- **Multi-body repulsive forces step in** and will generate explosion instead
- **Before ξ reaches inter-nucleon distances in ambient matter, it does so for clusters**
- **So, watching the clusters (which are very sensitive) is a better signal than ambient EOS**
- **We calculated main diagrams for different shapes and sizes of clusters**
- **We used universal Ising fluctuation potential,**
- **deformed because freeze out is away from critical line**
- **And get temperature dependence of effective triple and quartic couplings**
- **The results predict strong dip of clustering near TC**
- **Experimental data hint for TWO (?) correlated dips in TWO (very different) observables, kurtosis and tritium ratio**

Now at fixed $t=0.077$, $x_i=2$ fm but as a function of cluster size ρ

$\rho=2$ very strongly suppressed
Does it mean clusters have
Size $\rho=4$ fm?



Now at fixed $t=0.077$, $x_i=2$ fm but as a function of cluster size ρ

$\rho=2$ very strongly suppressed
Does it mean clusters have
Size $\rho=4$ fm?

No, it makes no sense since
then density is the same as ambient matter

